

ACQO Summer School  
November 29th-December 3rd 2004

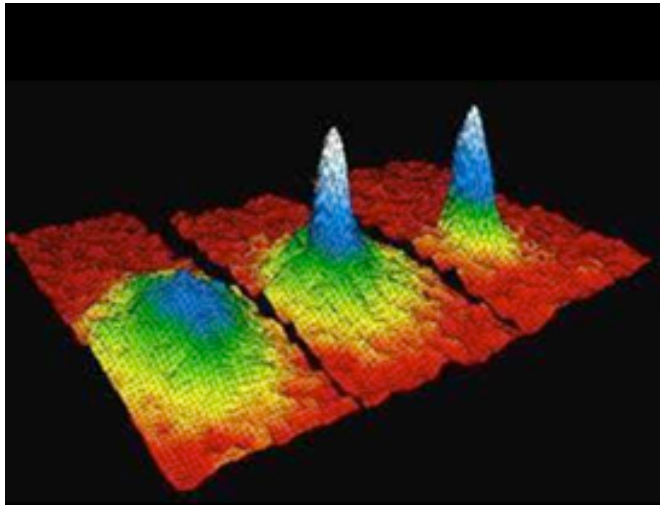
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Lecture 1

Experimental BEC: Back of the Envelope Calculations

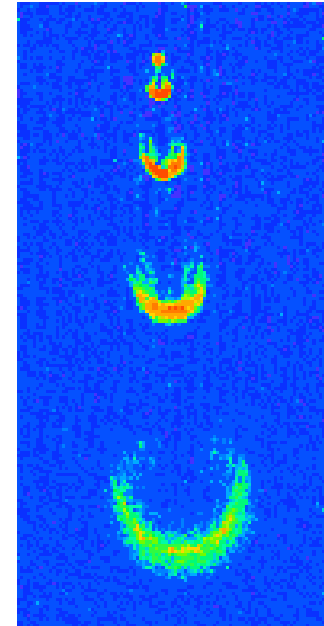
John Close  
ANU

## First BEC:JILA



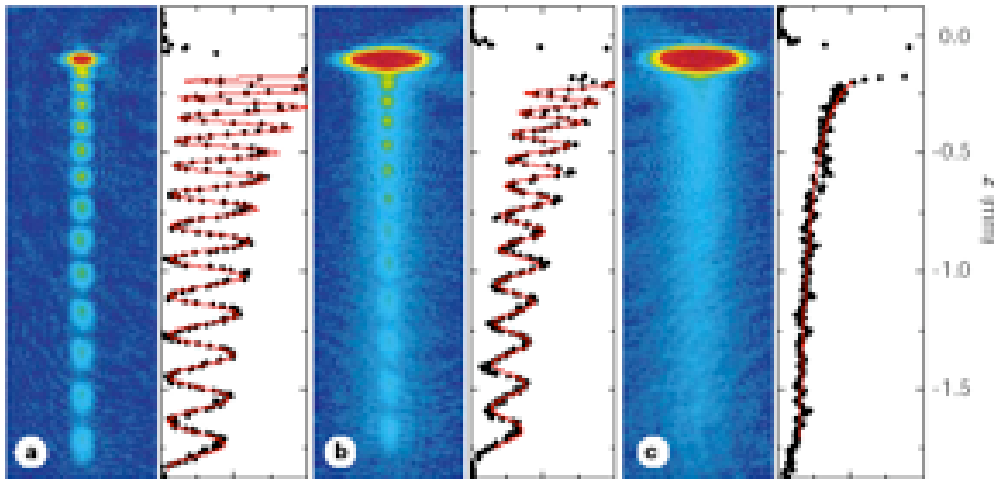
M.H. Anderson et al, Science, 269, 198 (1995).

## First Atom Laser:MIT



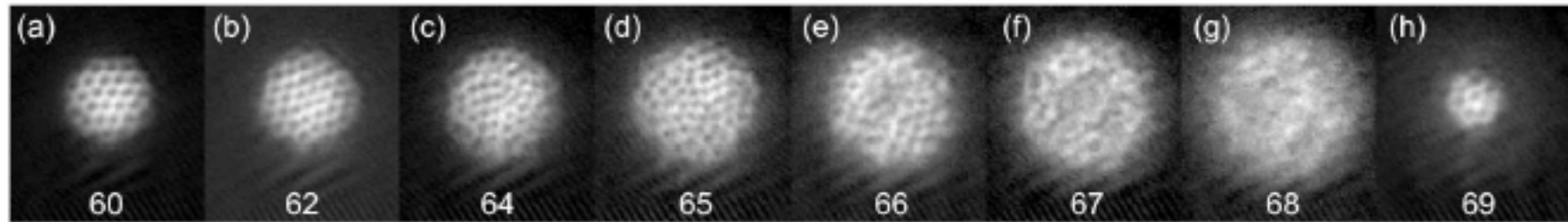
M. O. Mewes et al. ,PRL,**78**, 582 (1997).

## Coherence: MPI Munich



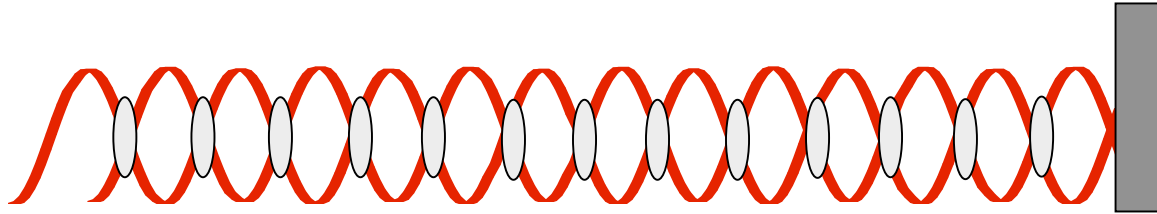
I.Bloch et al.,Nature 403,166, (2000).

## Vortices:ENS



Bretin et al., PRL, **92**, 050403-1 (2004)

## Quantum Fields and Coherence

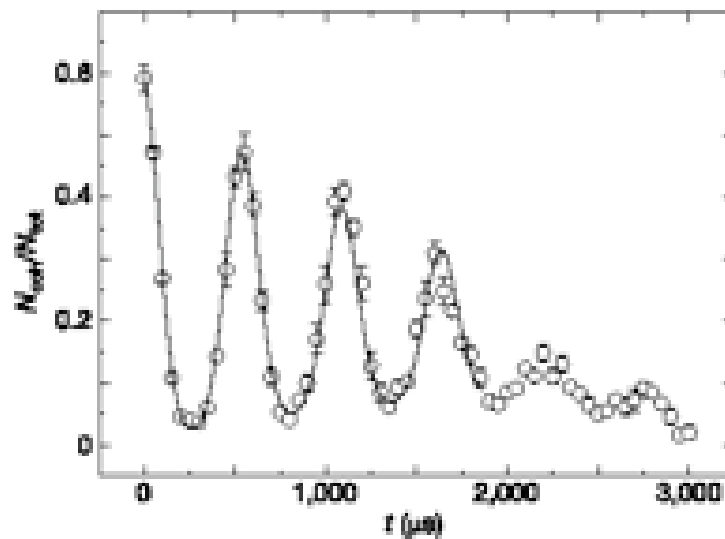
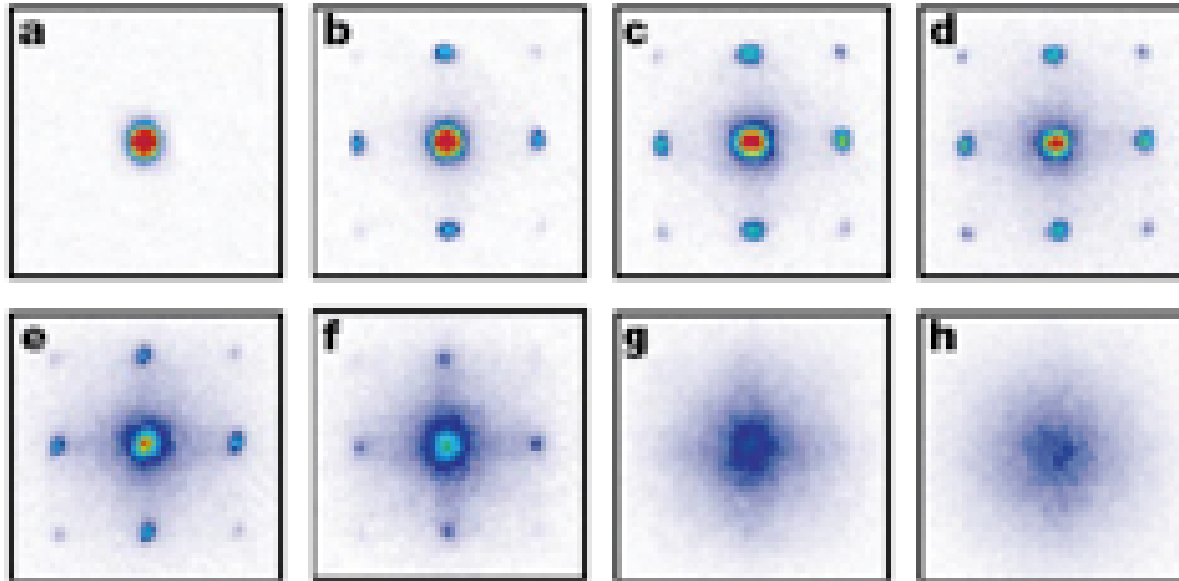


$$H = -J \sum_{i,j} \hat{a}_i^+ \hat{a}_j + \sum_i \varepsilon_i a_i^+ a_i + \frac{1}{2} U \sum_i a_i^+ a_i (a_i^+ a_i - 1)$$

$$|\Psi\rangle = |N\rangle \textit{ Fock state}$$

$$|\Psi\rangle = \exp(-|\alpha|^2) \sum_{n=0}^{\infty} \frac{\alpha^n}{\sqrt{n!}} |n\rangle \textit{ coherent state}$$

MPI Garching: Greiner et al,



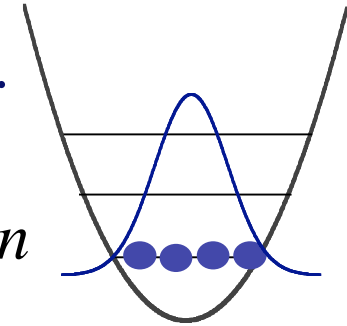
Nature 415, 40 (2002).  
Nature 419, 51 (2002).

$$\hat{\Psi}(\vec{r})$$

## What is a BEC ? Statistical Mechanics Definition

BEC is the macroscopic occupation of the ground state.

$$p(\varepsilon) = \frac{\exp(-\varepsilon/kT)}{\sum_{\text{all states}} \exp(-\varepsilon/kT)} \text{ Boltzmann}$$



### Bose Einstein

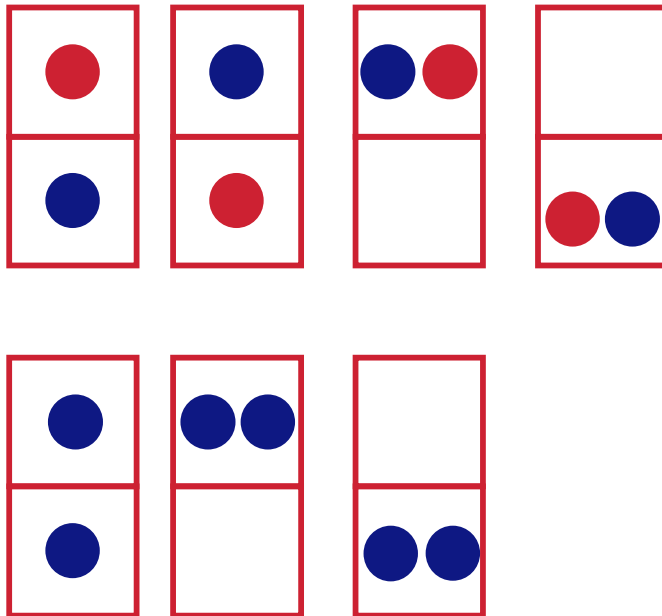
$$f(\varepsilon) = \frac{1}{\exp\left(\frac{\varepsilon - \mu}{kT}\right) - 1}$$

$$kT_c = \frac{2\pi\hbar^2}{m} \left( \frac{N}{2.612V} \right)^{\frac{2}{3}}$$

# What Drives BEC ?

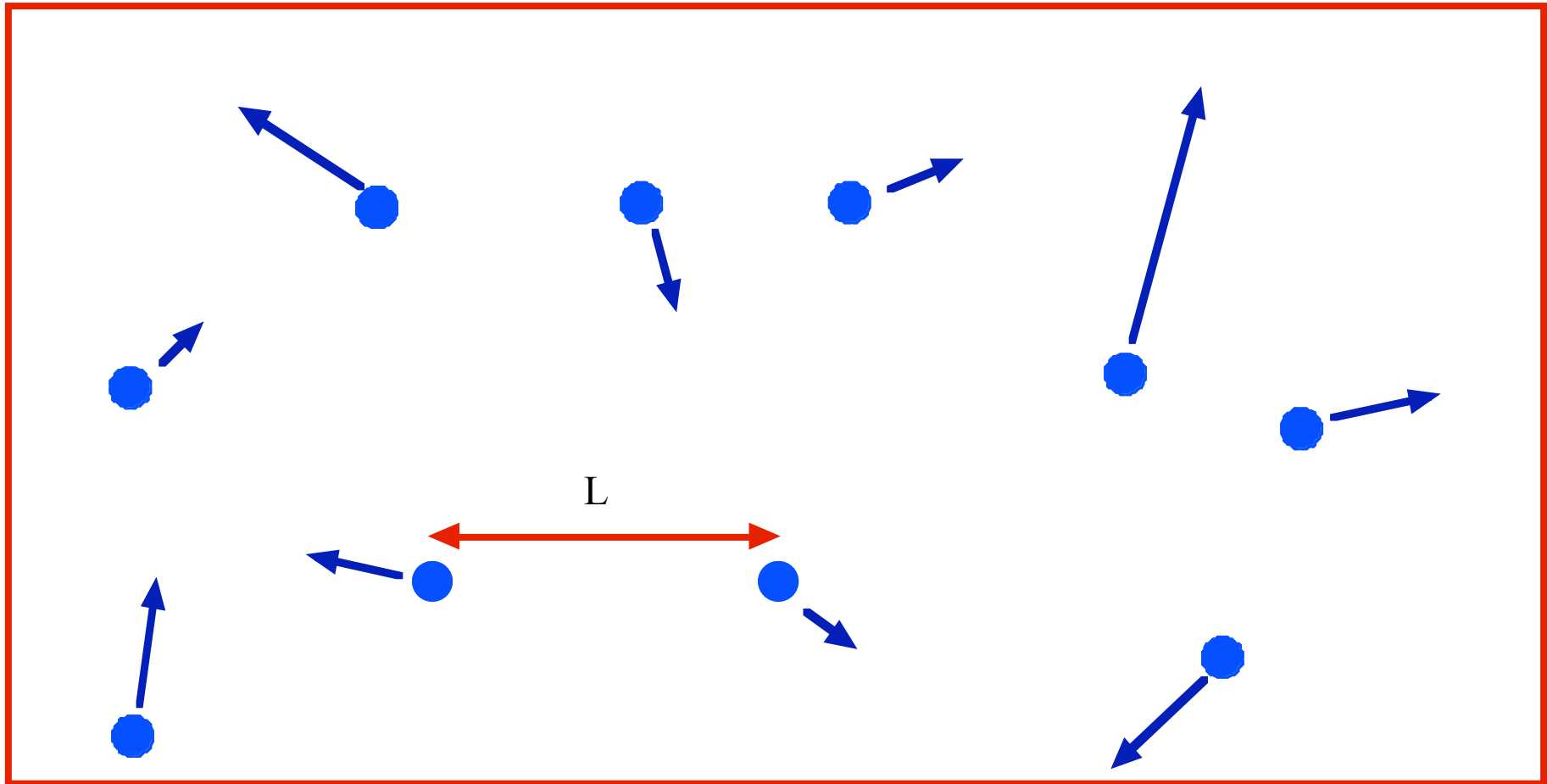
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## The Statistics of Indistinguishable Bosons



Ratio of the probability of  
ground state occupancy =  $4/3$

# A Classical Gas

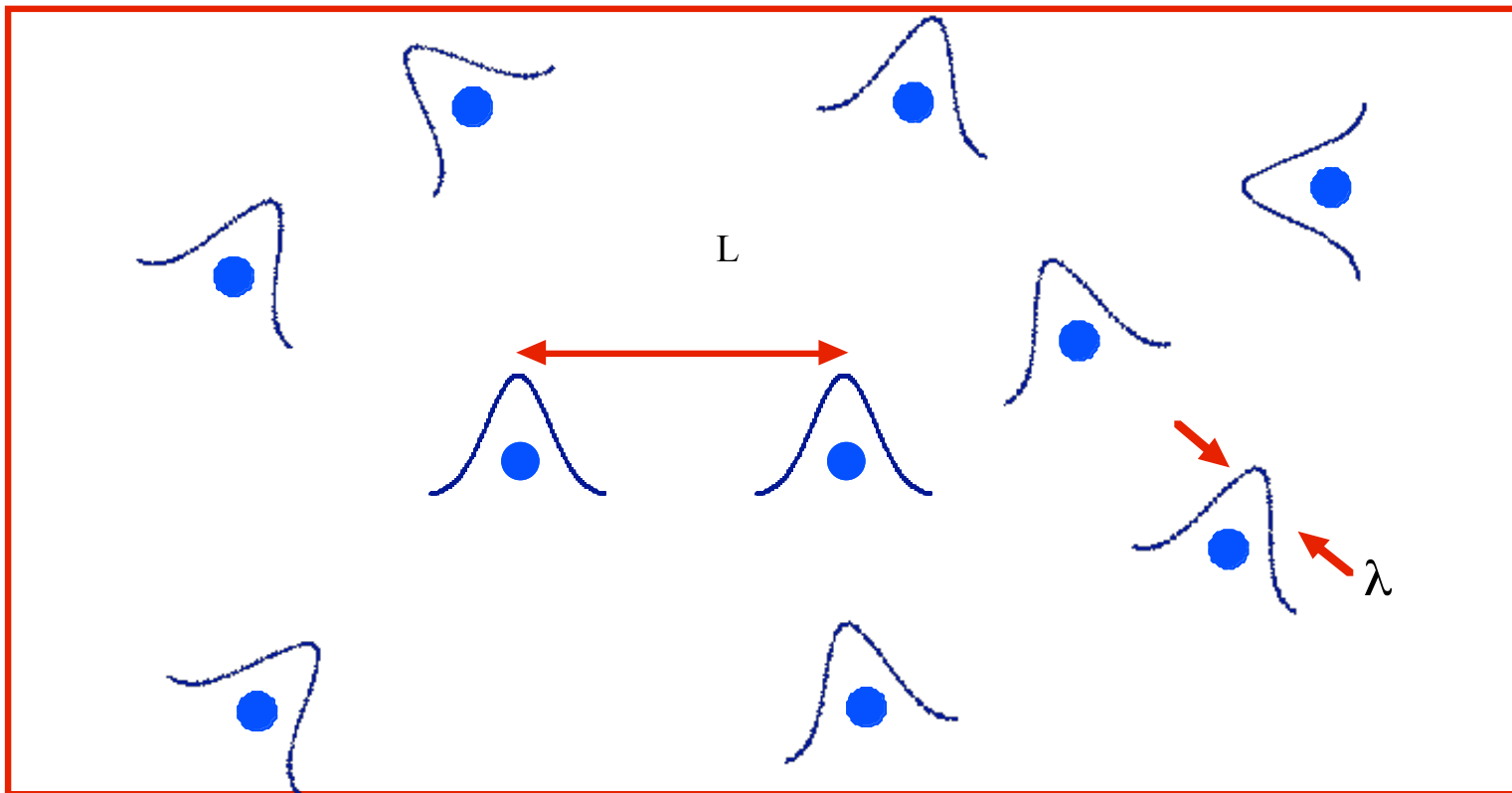


$$L = \frac{1}{\sqrt[3]{n}}$$

$$V = \sqrt{\frac{2kT}{m}}$$



# Gas With Delocalised Atoms



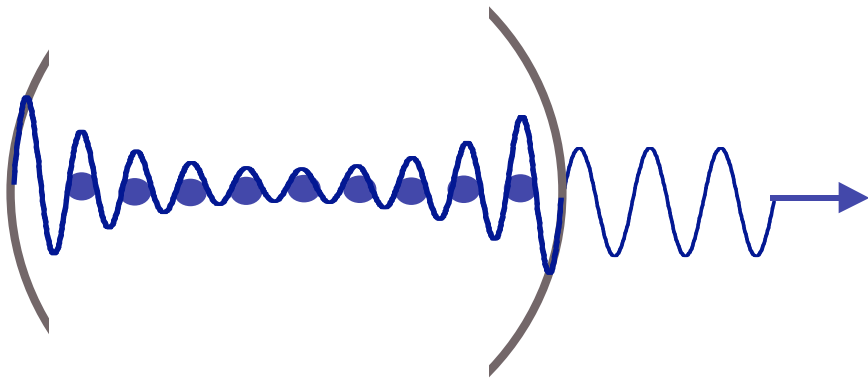
$$\lambda = \frac{\hbar}{p} = \hbar \sqrt{\frac{2\pi}{mkT}}$$

$$L = \frac{1}{\sqrt[3]{n}}$$

Phase space density

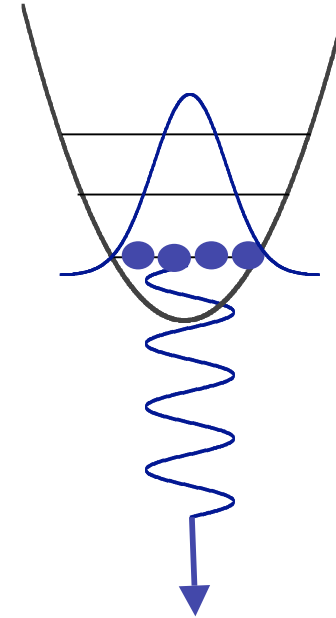
$$\rho = \left(\frac{\lambda}{L}\right)^3 = n \left(\frac{2\pi\hbar^2}{mkT}\right)^{\frac{3}{2}}$$

## The Optical Laser



- Macroscopic occupation of a cavity mode.
- Bright, coherent, polarised optical beam.

## The Atom Laser



- Macroscopic population of the ground state trap mode.
- Bright, coherent, polarised deBroglie matter beam.

# Atoms and Fields: How it all Works

## The simple story

$$\langle T \rangle = -\frac{1}{2} \langle V \rangle$$

$$a_0 = \frac{4\pi\epsilon_0\hbar^2}{me^2} = 0.5 \text{ \AA} \quad \text{Size of an atom}$$

$$\left\langle \frac{p^2}{2m} \right\rangle = \frac{e^2}{8\pi\epsilon_0} \left\langle \frac{1}{a_0} \right\rangle$$

$$pa_0 \approx \hbar$$

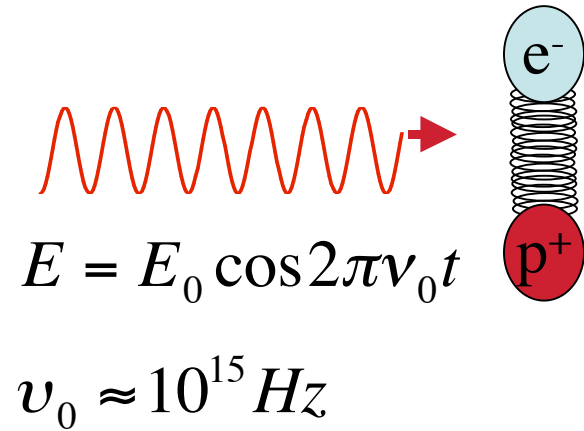
$$\Delta E = \frac{\hbar^2}{8ma_0^2} \approx 3eV$$

Atoms absorb  
In the visible range !

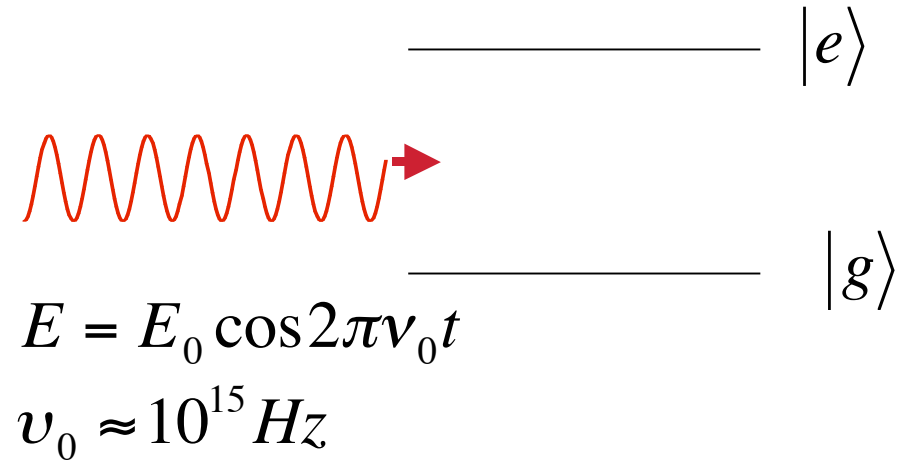
$$\Delta\nu \approx 5 \times 10^{14} \text{ Hz}$$

# Simple models that will get us a long way

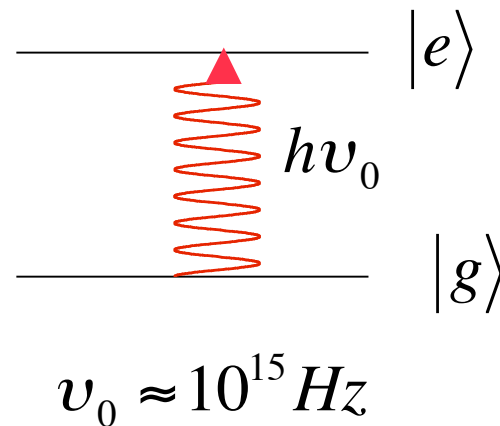
## Classical



## Semi-classical

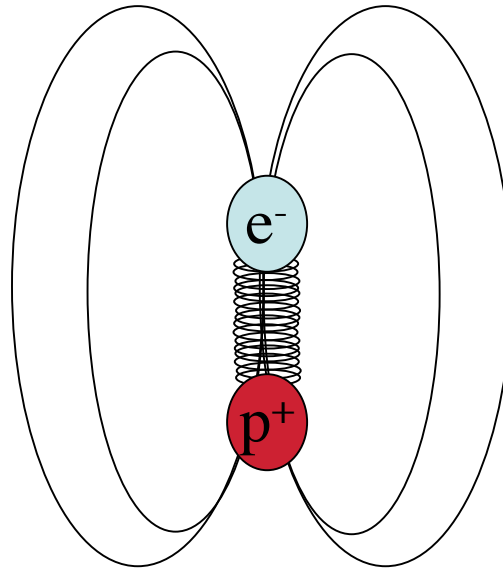


## Quantum



## How Long Does an Atom Stay in the Excited State ?

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$$\text{power radiated by a classical dipole} = \frac{\wp^2 \omega^4}{12\pi\epsilon_0 c^3}$$

$$\text{spontaneous emission rate} = \Gamma = \frac{\wp^2 \omega^4}{12\pi\epsilon_0 c^3} \frac{1}{\hbar\omega}$$

$$\Gamma = \frac{\wp^2 \omega^3}{12\pi\hbar\epsilon_0 c}$$

$$\Gamma = \frac{\wp^2 \omega^3}{3\pi\hbar\epsilon_0 c^3}$$

## So how fast do atoms scatter photons ?

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$$\wp \approx ea_0 \approx 8 \times 10^{-29} \text{ Cm}$$

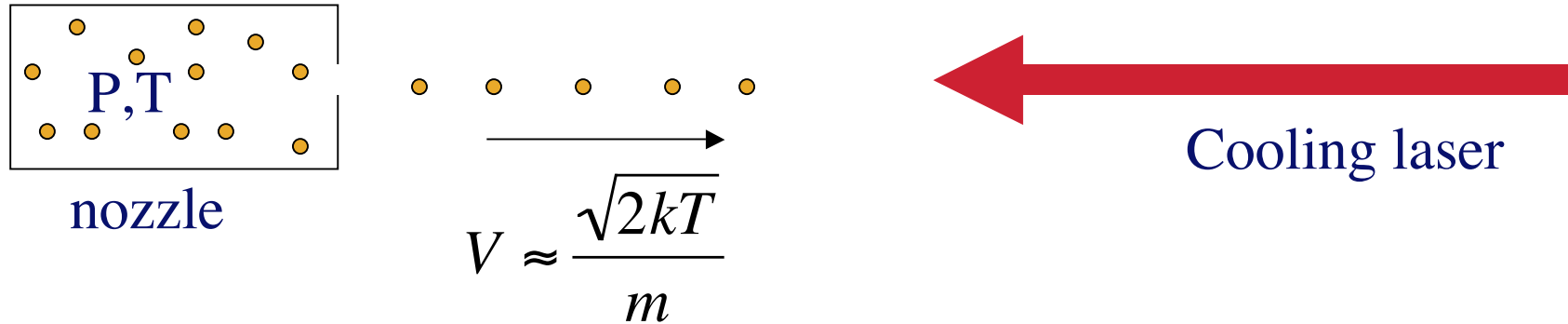
$$\omega \approx 2\pi \times 5 \times 10^{14} \text{ Hz}$$

$$\Gamma \approx 10 \text{ MHz}$$

### What have we found out so far ?

- 1) Atoms are roughly  $5 \times 10^{-11}$  metres in size.
- 2) Atoms absorb light in the visible spectrum ( $\nu \approx 10^{15} \text{ Hz}$ )
- 3) Atoms can scatter roughly  $10^7$  photons per second.

# How long does it take to stop a room temperature atom ?



$$a = \frac{F}{m} = \frac{\Gamma \hbar k}{m} = \frac{\Gamma h \nu}{mc} \approx 2 \times 10^5 \text{ ms}^{-2}$$

$$t_{stop} = \frac{V}{a} = \frac{1}{a} \underbrace{\sqrt{\frac{2kT}{m}}}_{300 \text{ m/s}} \approx 10^{-3} \text{ s}$$

$$x_{stop} \approx 1 \text{ m}$$

## What Intensity do we need in our laser to achieve this cooling ?

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$$I_{\text{saturation}} = 10^7 \text{ photons} / \underbrace{\text{photon scattering cross section} / \text{second}}_{\sigma}$$

$$\sigma \approx \pi a_0^2 \approx 10^{-21} \text{ m}^2 ?$$

$$I_{\text{saturation}} = 10^{28} \text{ photons} / \text{m}^2 / \text{s}$$

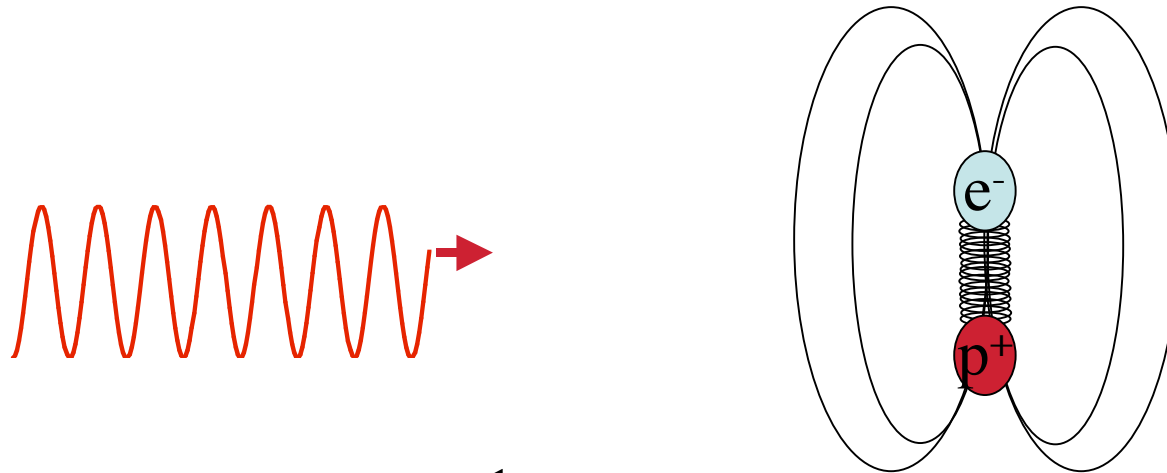
$$1 \text{ visible photon} \approx 3 \times 10^{-19} \text{ J} \approx 2 \text{ eV}$$

$$1 \text{ Watt of visible photons} \approx 3 \times 10^{18} \text{ photons} / \text{second}$$

$$I_{\text{saturation}} = 10^{28} \text{ photons} / \text{m}^2 / \text{s} = 300 \text{ kW} / \text{cm}^2$$



## Calculate the scattering cross section from a driven dipole



$$m \ddot{x} + \alpha \dot{x} + \frac{1}{2} m \omega^2 x^2 = E_0 e^{i\omega t}$$

Solve the equation of motion above and calculate the power dissipated by the dipole in terms of  $\alpha$ . Set this equal to the power radiated by the dipole to determine  $\alpha$ . Then use:

$$\sigma = \frac{P_{\text{radiated}}}{I_{\text{incident}}} = \frac{2}{c \epsilon_0 E_0^2} P_{\text{radiated}} = \frac{3\lambda^2}{2\pi} \approx 10^{-13} m^2$$

What Intensity do we need from our laser to achieve this cooling ?

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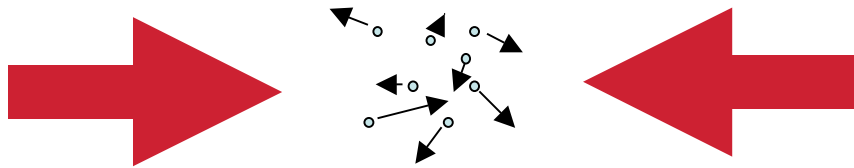
$$I_{\text{saturation}} = 10^7 \text{ photons} / \underbrace{\text{photon scattering cross section} / \text{second}}_{\sigma}$$

$$\sigma = \frac{3\lambda^2}{2\pi} \approx 10^{-13} \text{ m}^2 ?$$

$$I_{\text{saturation}} = 10^{20} \text{ photons} / \text{m}^2 / \text{s}$$

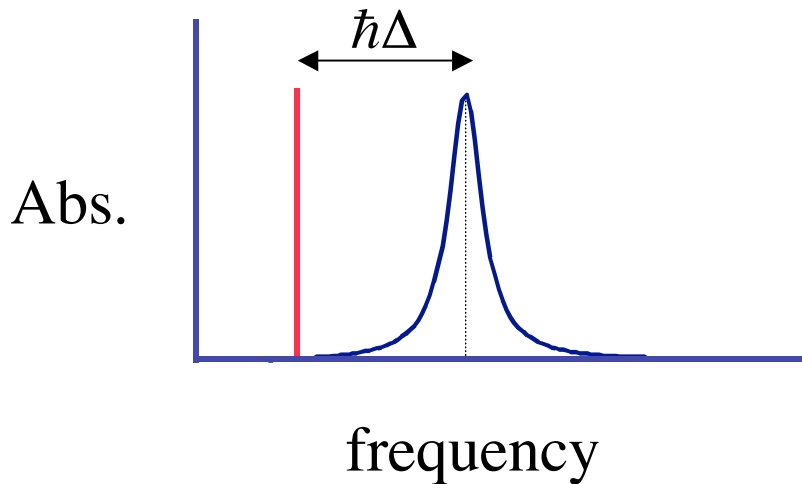
$$I_{\text{saturation}} = 10^{20} \text{ photons} / \text{m}^2 / \text{s} = 3 \text{ mW} / \text{cm}^2$$

# What is the limit to this kind of cooling ?



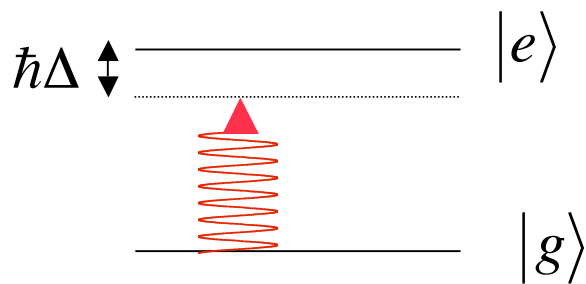
$$Force = F(V, k, \Delta)$$

$$F = \hbar \frac{\Delta}{\Gamma} k^2 V$$



$$\left( \frac{dE}{dt} \right)_{cool} = FV = \frac{2\hbar k^2}{M} E$$

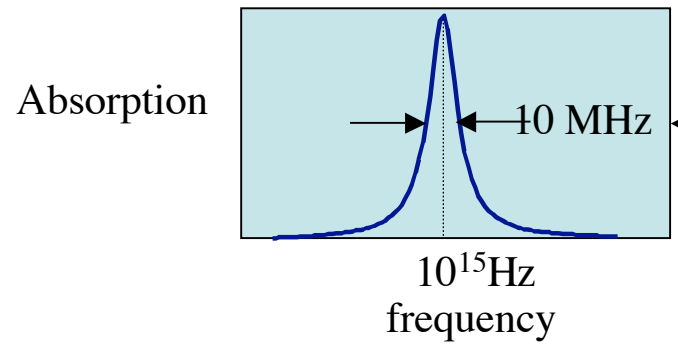
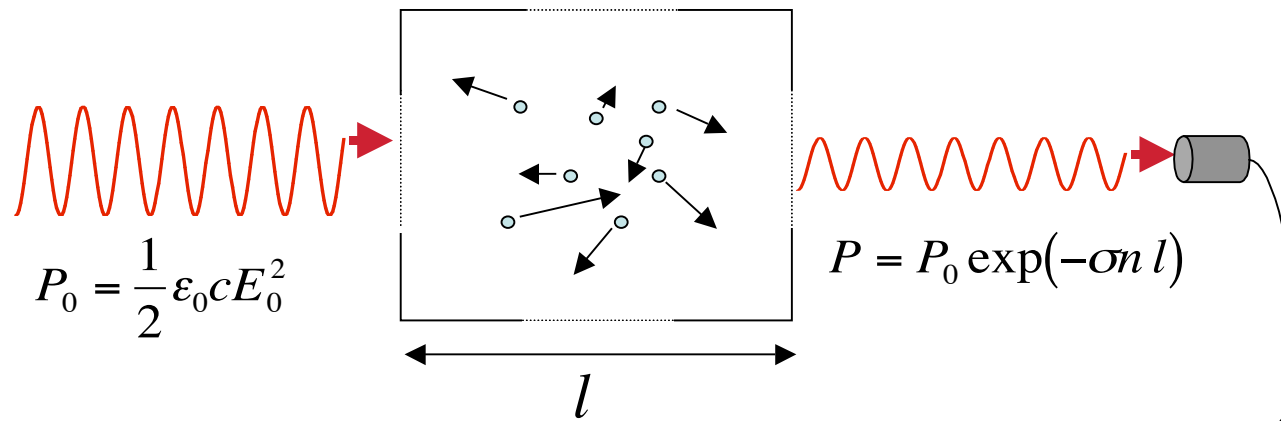
$$\left( \frac{dE}{dt} \right)_{heat} = \frac{1}{2M} \frac{d\langle p^2 \rangle}{dt} = \frac{\Gamma(\hbar k)^2}{2M}$$



$$E_{\min} = \frac{\hbar\Gamma}{4} \quad T_{\min} = \frac{\hbar\Gamma}{4k_b}$$

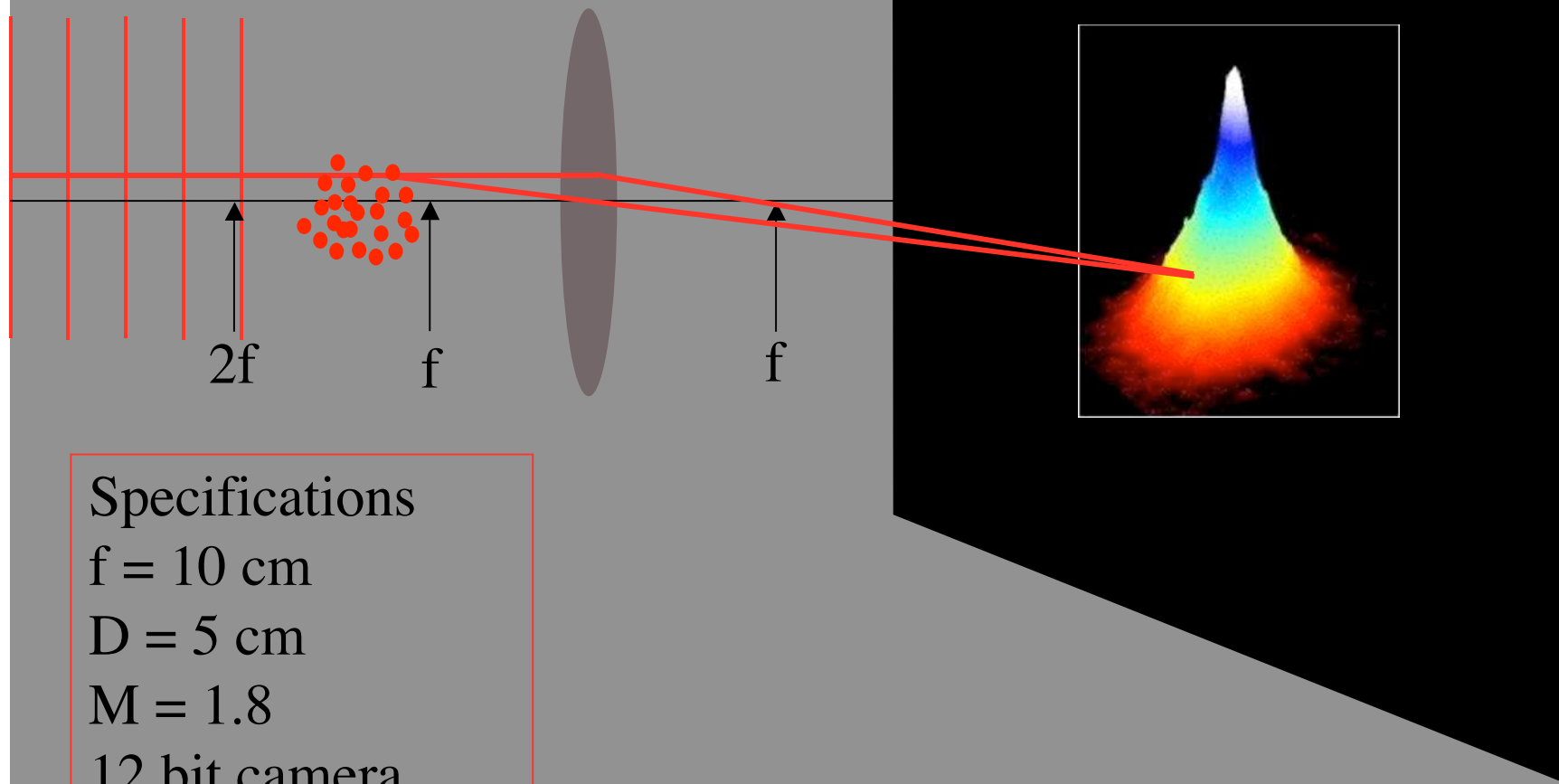
$$T_{\min} = \frac{\hbar\Gamma}{2k_b} \approx 100\mu K \text{ for } \Gamma = 10 \text{ MHz}$$

# Absorption Imaging



$$n = \frac{\ln P - \ln P_0}{\sigma l}$$

## Absorption Imaging of Condensates



### Specifications

$f = 10$  cm

$D = 5$  cm

$M = 1.8$

12 bit camera

$9\mu\text{m} \times 9\mu\text{m}$  pixels

## We have established the following

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### Atoms:

- are on the order of  $10^{-10}$  m in size.
- absorb in the visible range.
- can scatter roughly  $10^7$  photons per second.
- have a photon scattering cross section of roughly  $10^{-13}\text{m}^2$
- can be laser cooled to a temperature on the order of  $100\ \mu\text{K}$  in a time on the order of milliseconds by laser beams with intensities on the order of  $\text{mW}/\text{cm}^2$ . The cooling laser linewidth should be  $1\text{MHz}$  or less.