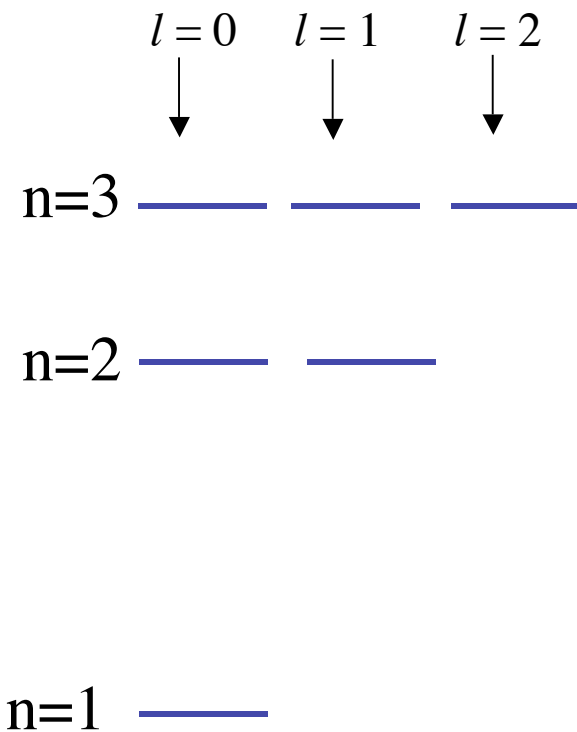


Real Atoms

$$H = \frac{\hat{p}^2}{2m_e} - \frac{e^2}{4\pi\epsilon_0 r}$$



Good quantum numbers

n, l, m_l, s, m_s

Eigenstates

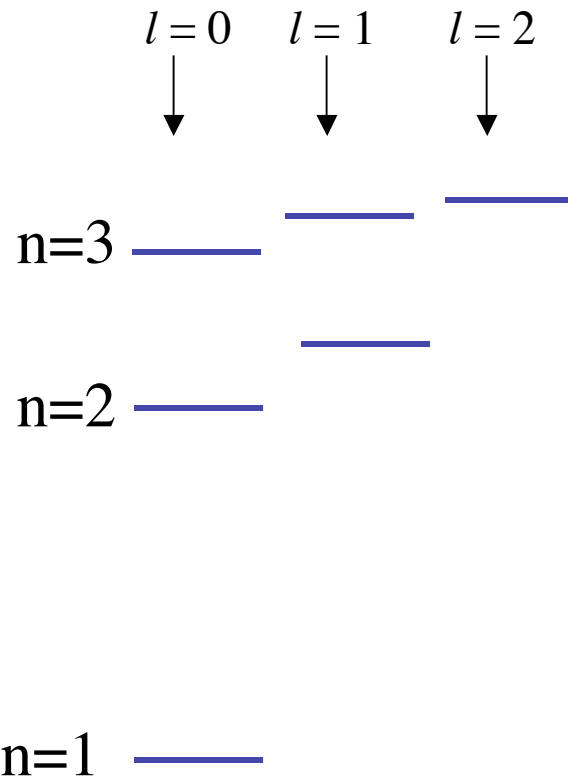
$$\Psi_{n,l,m}(\vec{r}) = R_{n,l}(r)Y_l^m(\theta, \phi)$$

Energy eigenvalues

$$E_n = -\frac{1}{2}mc^2 \frac{\alpha}{n^2} = \frac{-13.6eV}{n^2}$$

Alkali Atoms

$$H = \frac{\hat{p}^2}{2m_e} + U(r) - \frac{Ze^2}{4\pi\epsilon_0 r}$$



Good quantum numbers

n, l, m_l, s, m_s

Eigenstates

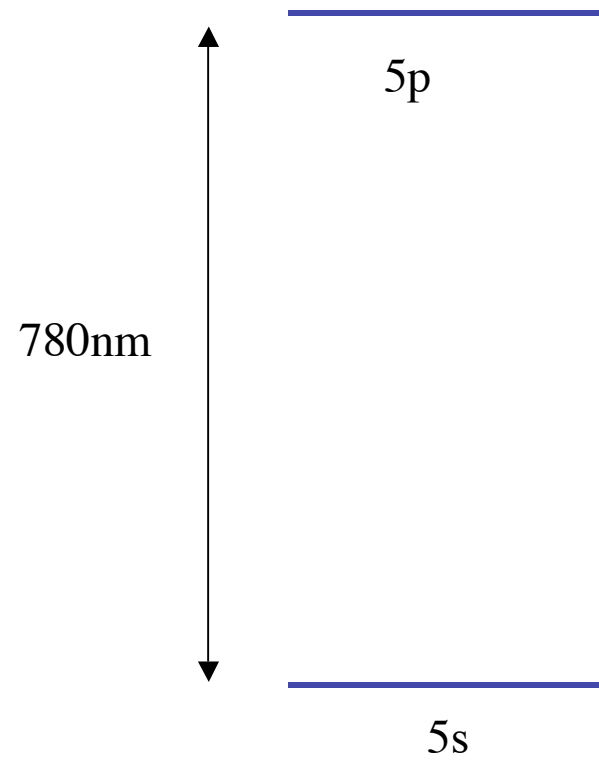
$$\Psi_{n,l,m}(\vec{r}) = R_{n,l}(r)Y_l^m(\theta, \phi)$$

Energy eigenvalues

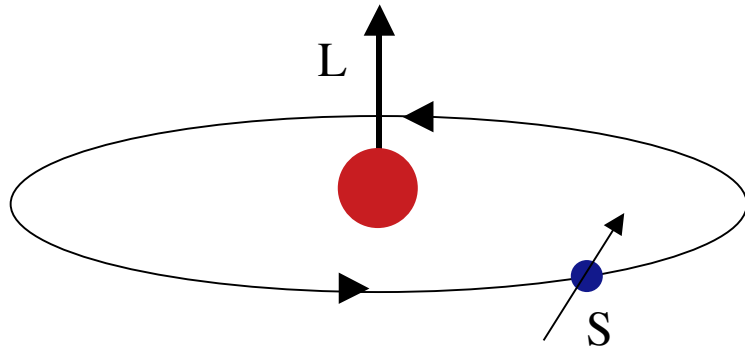
For quantum chemists to calculate

Rubidium

$$H = \frac{\hat{p}^2}{2m_e} + U(r) - \frac{Ze^2}{4\pi\epsilon_0 r}$$



Spin Orbit Coupling



$$H = -\vec{\mu} \cdot \vec{B} \quad H_{so} \propto \vec{S} \cdot \vec{L}$$

$$\vec{J} = \vec{L} + \vec{S}$$

$$J^2 |J, M_J, L, S\rangle = J(J+1)\hbar^2 |J, M_J, L, S\rangle$$

$$J_z |J, M_J, L, S\rangle = M_J \hbar |J, M_J, L, S\rangle$$

$$L \cdot S = \frac{1}{2}(J^2 - L^2 - S^2)$$

$$H_{so} \propto J^2 - L^2 - S^2$$

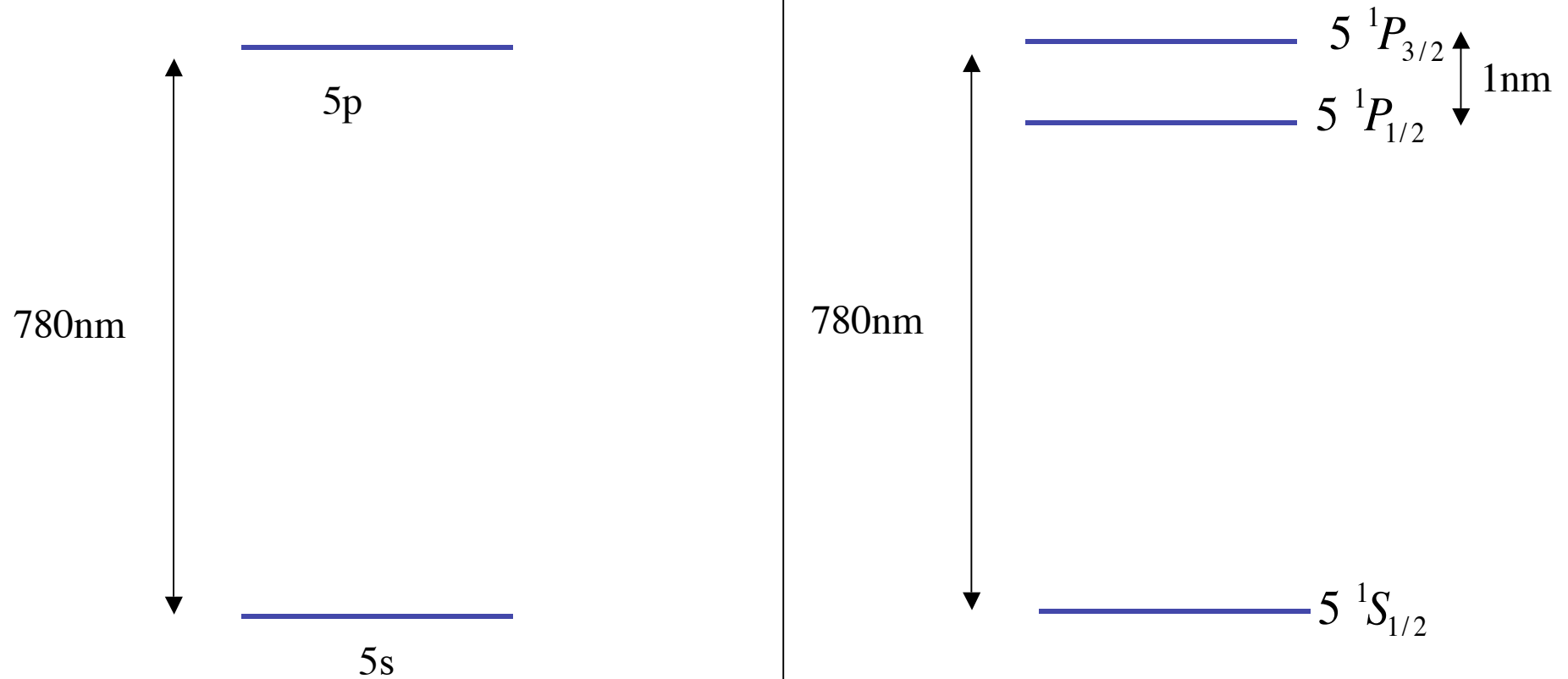
Relativistic correction

$$E_k = \sqrt{c^2 p^2 + m_e^2 c^4}$$

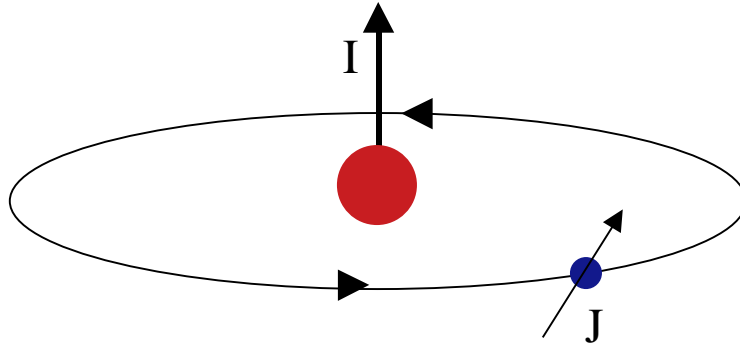
$$E_k = m_e c^2 + \frac{p^2}{2m} - \frac{p^4}{8m^3 c^2} + \dots$$

Fine Structure

$$H \approx \underbrace{\frac{\hat{p}^2}{2m_e} - \frac{Ze^2}{4\pi\epsilon_0 r}}_{\text{hydrogen like}} + \underbrace{\frac{e^2}{4\pi\epsilon_0} \sum_{i=1}^{N-1} \int d^3r_i \frac{|\Psi(r_i)|^2}{|\vec{r}_i - \vec{r}|}}_{\text{core electrons}} + \underbrace{\frac{e^2}{8\pi\epsilon_0 m_e^2 c^2} \frac{1}{r^3} \vec{S} \cdot \vec{L}}_{\text{spin orbit}} - \underbrace{\frac{p^4}{8m_e^3 c^2}}_{\text{relativistic}}$$



Hyperfine Interaction



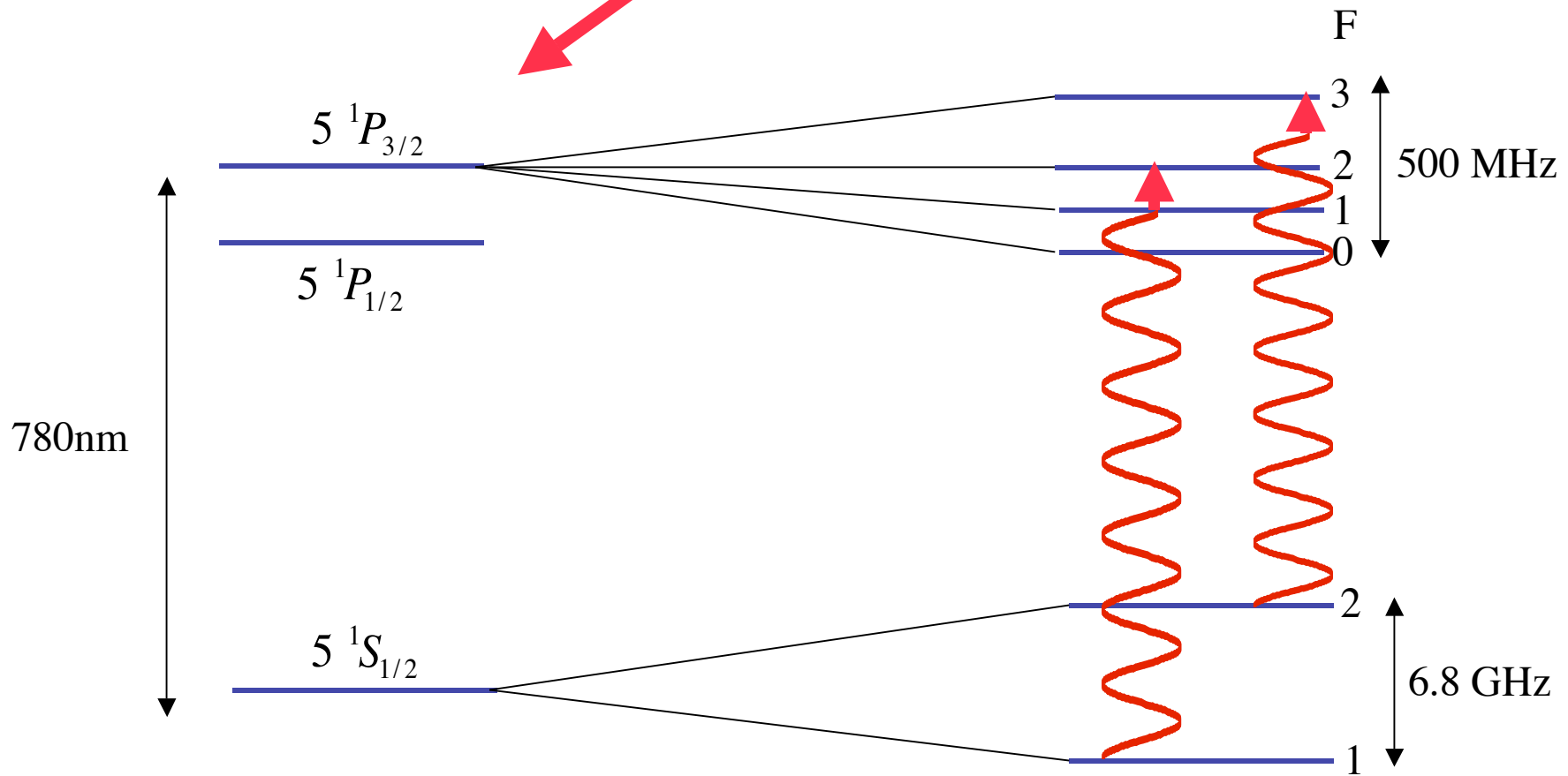
$$H_{HF} = A(J)J \cdot I$$

$$F = J + I$$

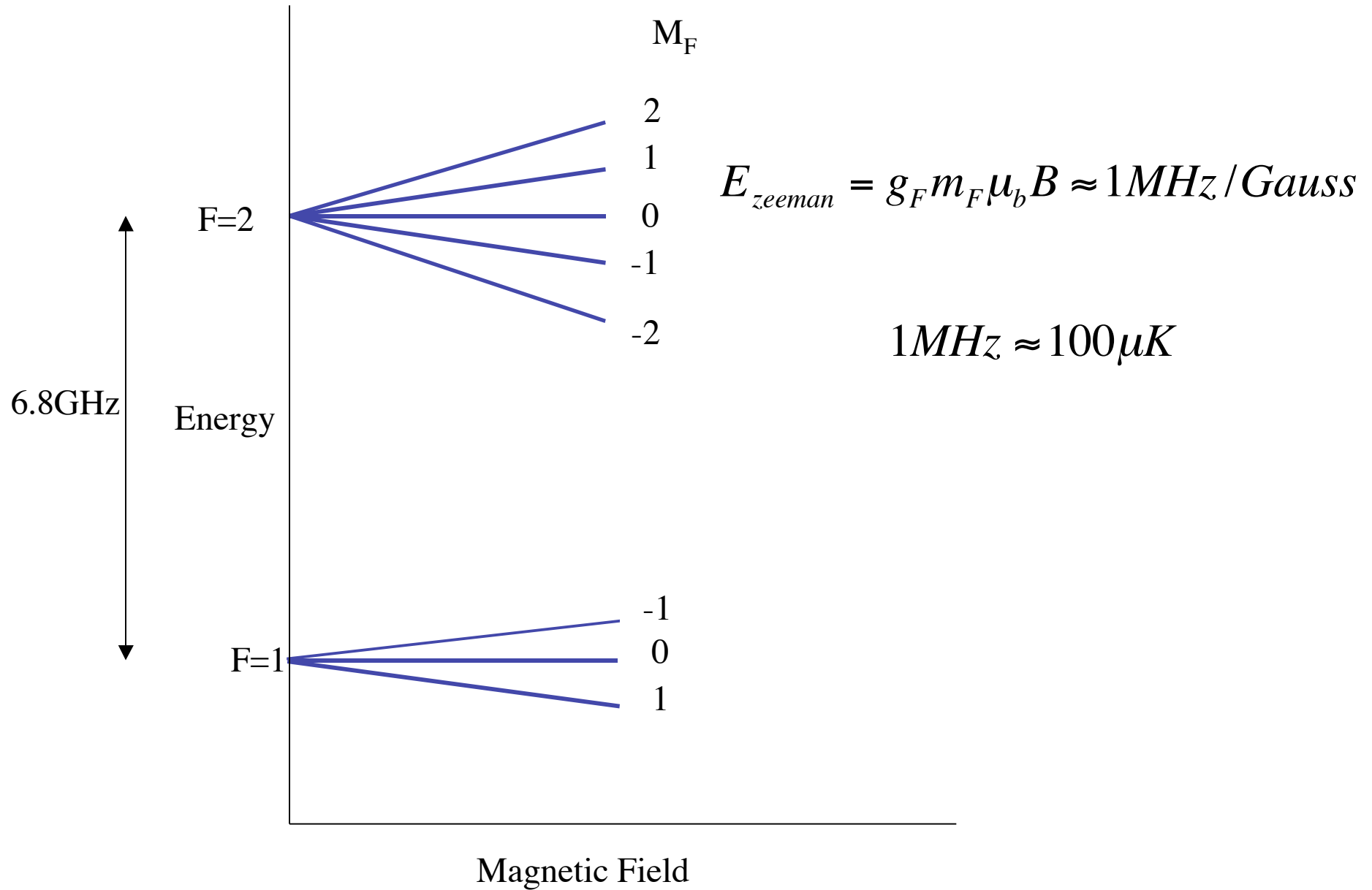
$$H_{HF} = A(J)(F^2 - J^2 - I^2)$$

$$H = \underbrace{\frac{\hat{p}^2}{2m_e} - \frac{Ze^2}{4\pi\epsilon_0 r}}_{\text{hydrogen like}} + \underbrace{\frac{e^2}{4\pi\epsilon_0} \sum_{i=1}^{N-1} \int d^3r_i \frac{|\Psi(r_i)|^2}{|\vec{r}_i - \vec{r}|}}_{\text{core electrons}} + \underbrace{\frac{e^2}{8\pi\epsilon_0 m_e^2 c^2} \frac{1}{r^3} \vec{S} \cdot \vec{L}}_{\text{spin orbit}} - \underbrace{\frac{p^4}{8m_e^3 c^2}}_{\text{relativistic}} + \underbrace{\xi \vec{J} \cdot \vec{I}}_{\text{hyperfine}}$$

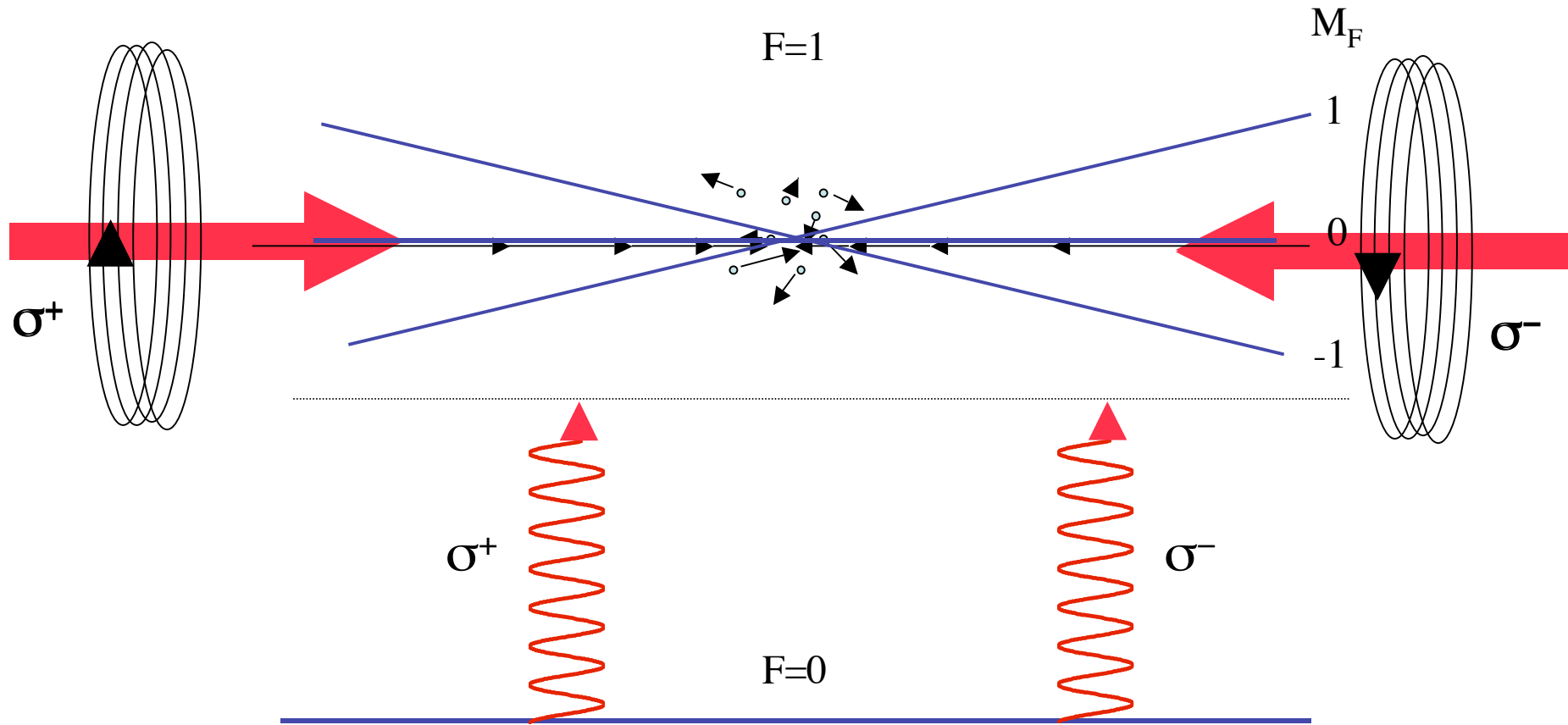
Including Hyperfine (^{87}Rb : $I=3/2$)



The Zeeman Effect and Magnetic Trapping



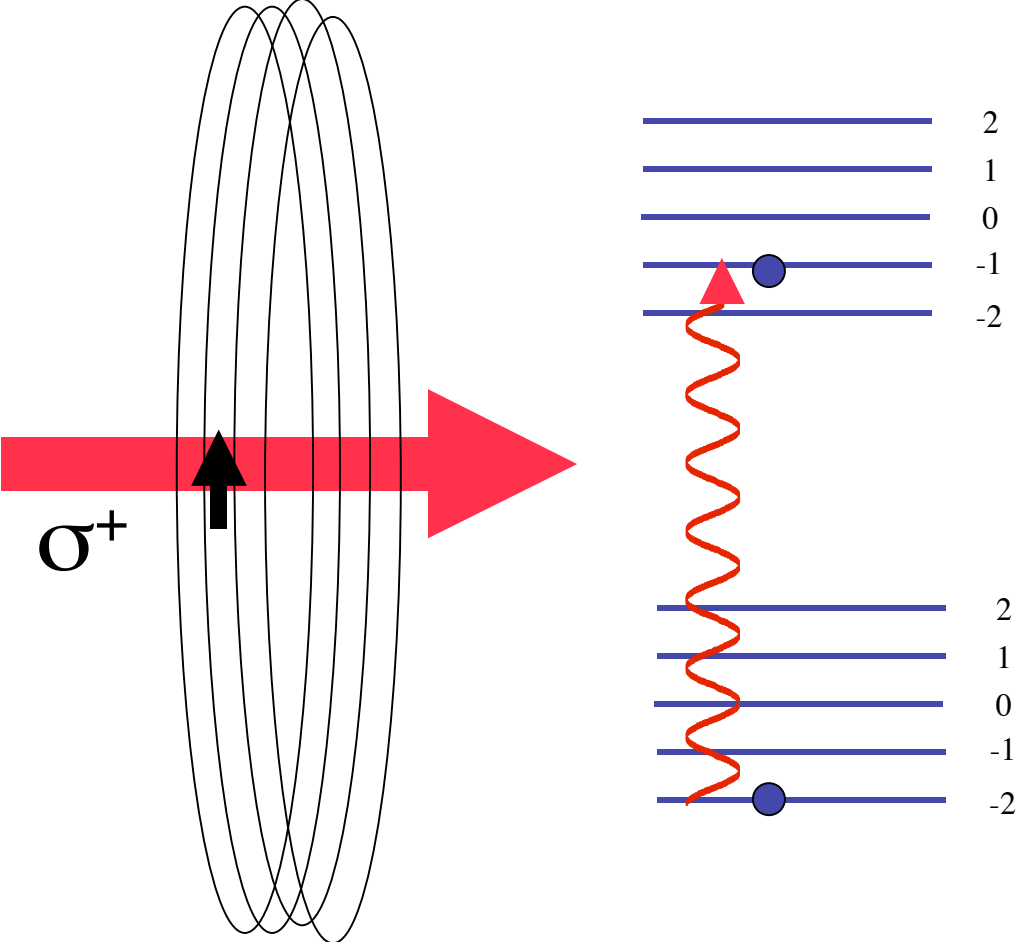
The Magneto Optic Trap



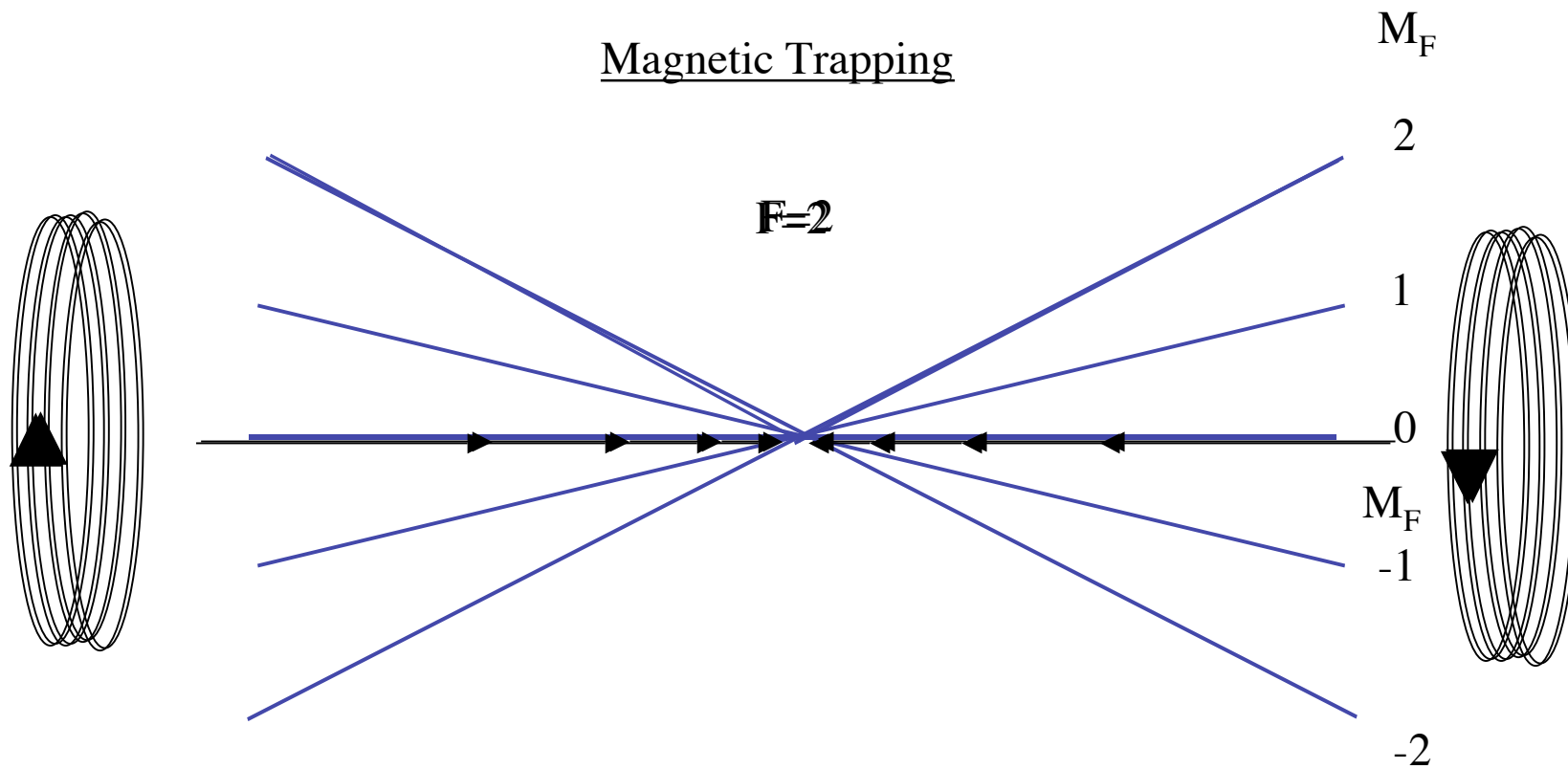
Typical Numbers

$$n=10^{10} \text{ cm}^{-3}$$
$$T=200 \text{ } \mu\text{K}$$
$$\rho=10^{-6}$$

Optical Pumping

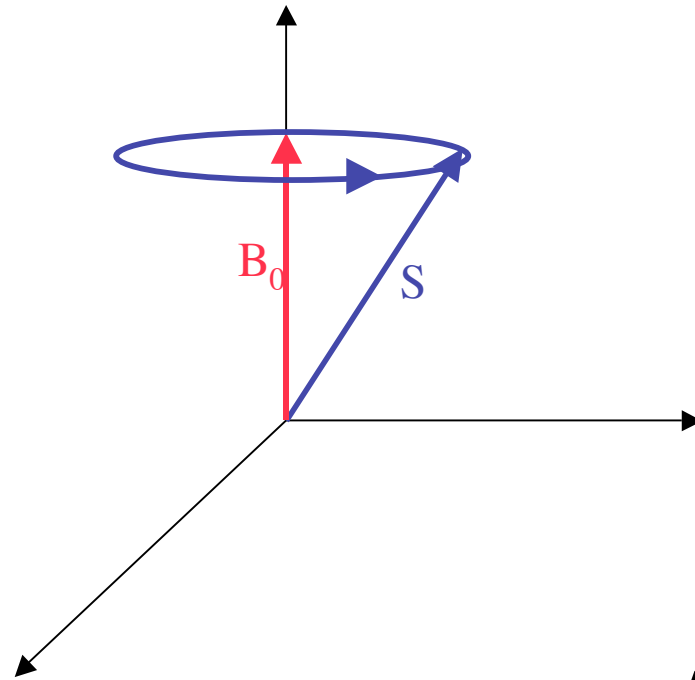


Magnetic Trapping



$$U(\vec{r}) = g_F m_F \mu_b |B(\vec{r})| \approx 1 \text{ MHz / Gauss}$$

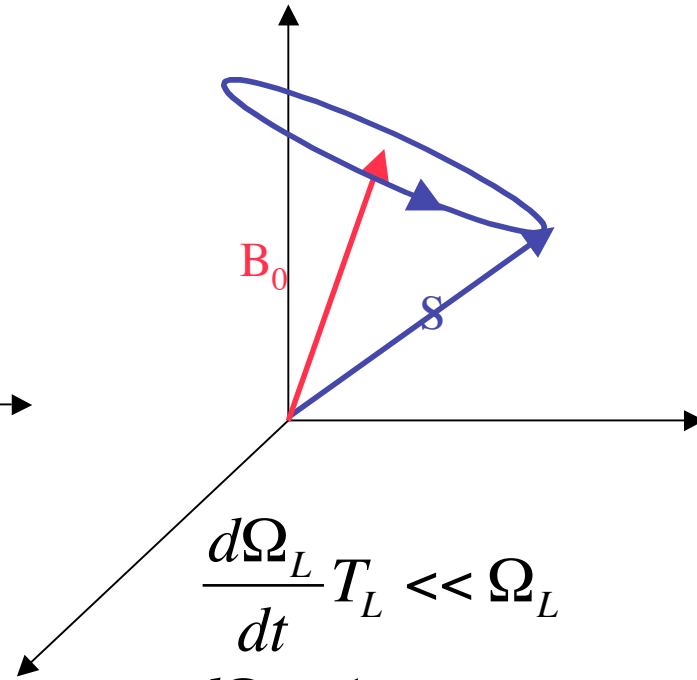
Classical Spin in a field



$$\frac{d\vec{S}}{dt} = \vec{\mu} \times \vec{B} = \gamma \vec{S} \times \vec{B}$$

The spin precesses around \mathbf{B}_0 at the Larmour precession frequency

$$\Omega_L = \gamma B_0 = g_F m_F \mu_b B \approx 1 \text{ MHz/Gauss}$$



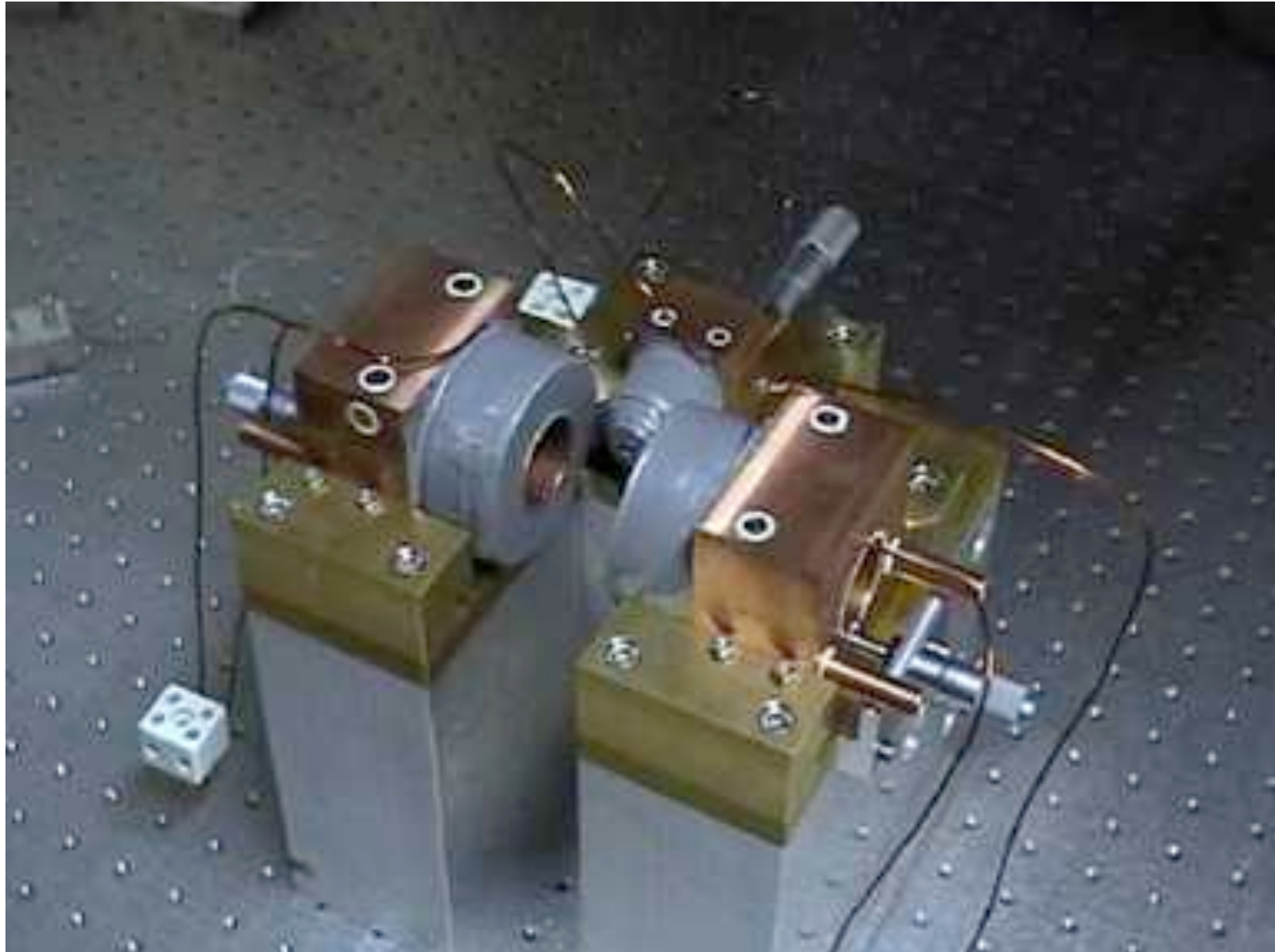
$$\frac{d\Omega_L}{dt} T_L \ll \Omega_L$$

$$\frac{d\Omega_L}{dt} \frac{1}{\Omega_L} \ll \Omega_L$$

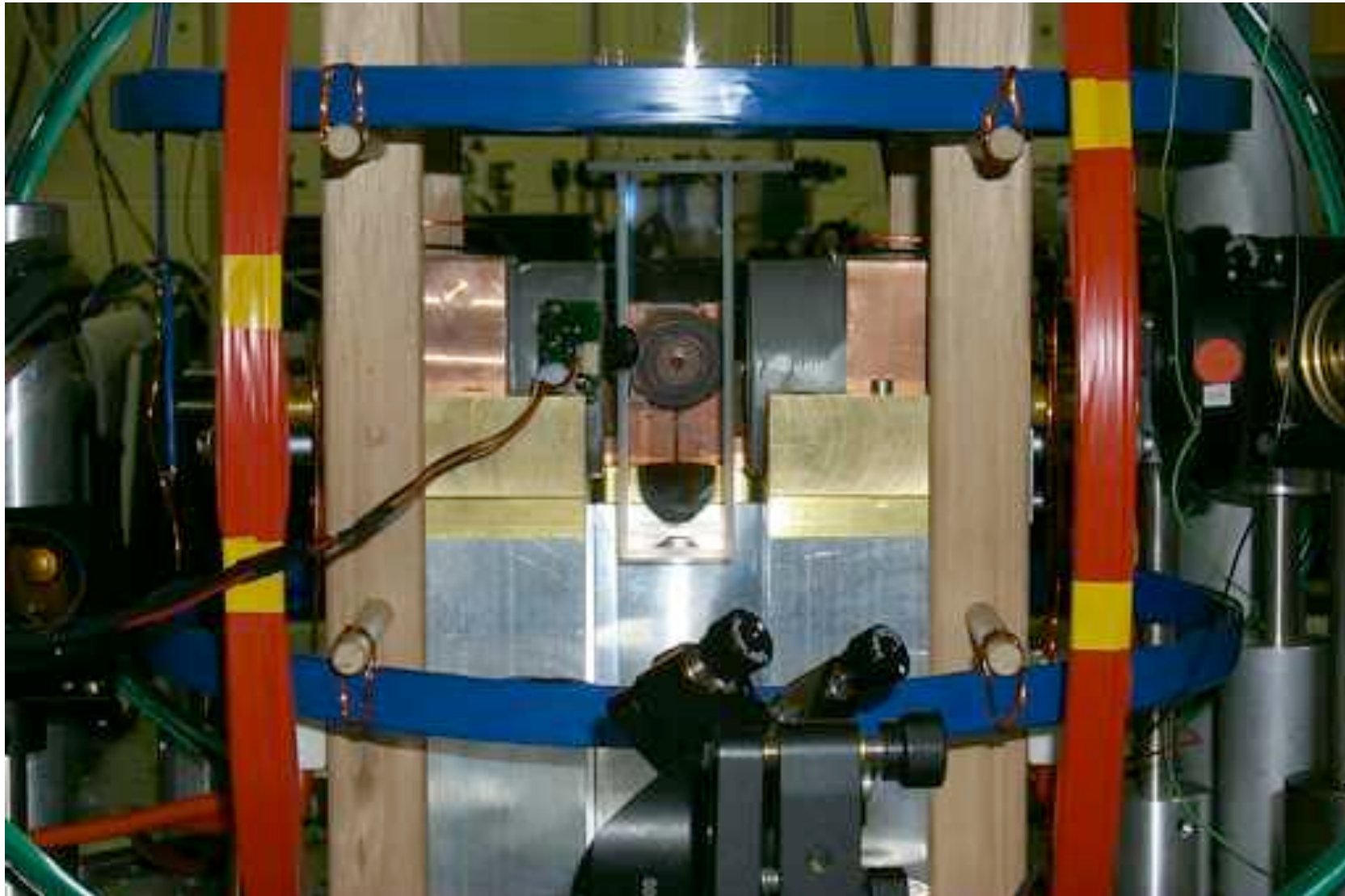
$$\frac{d\Omega_L}{dt} \ll \Omega_L^2$$

$$(V \cdot \nabla) B \ll \gamma B^2$$

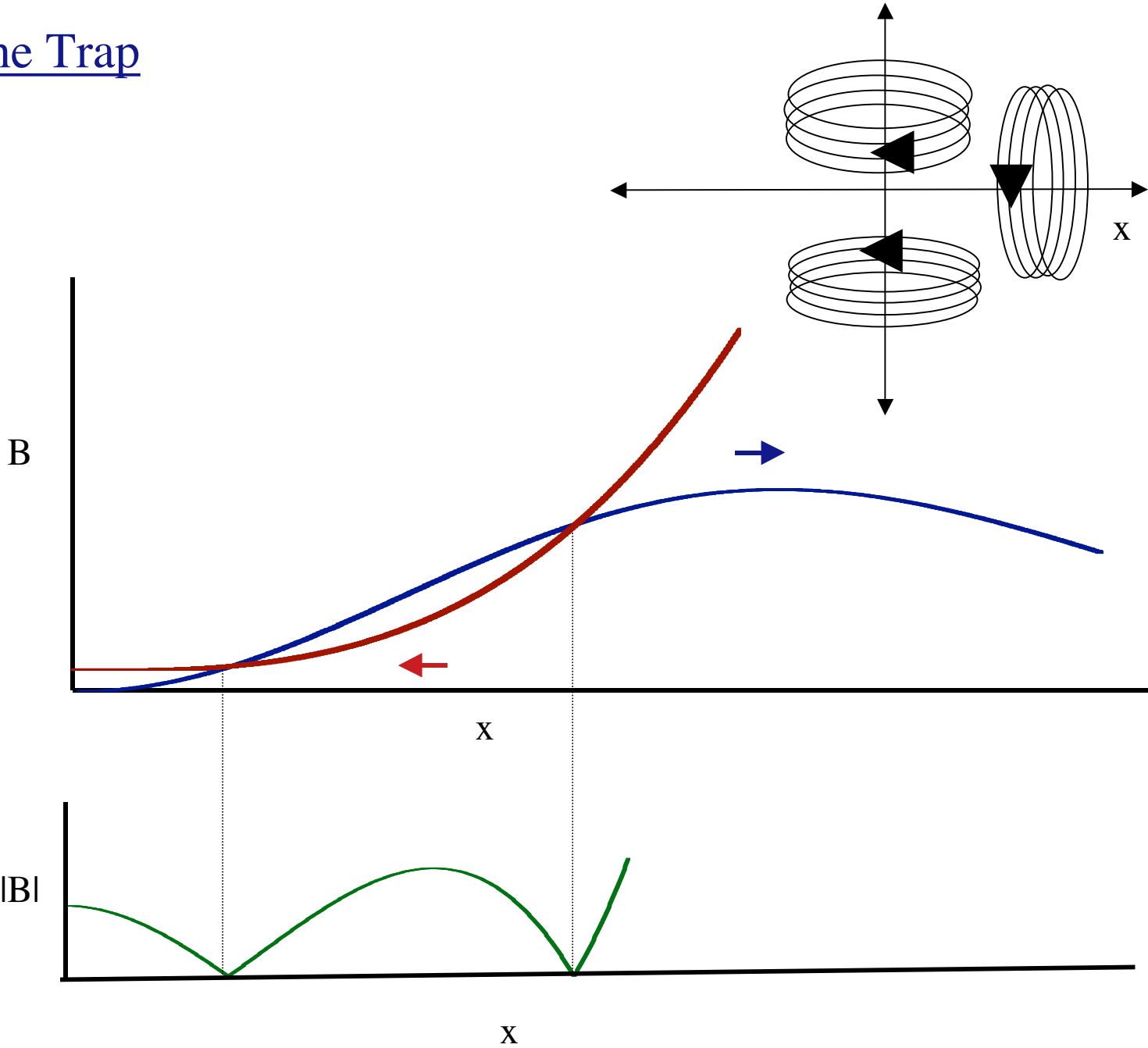
The Magnetic Trap



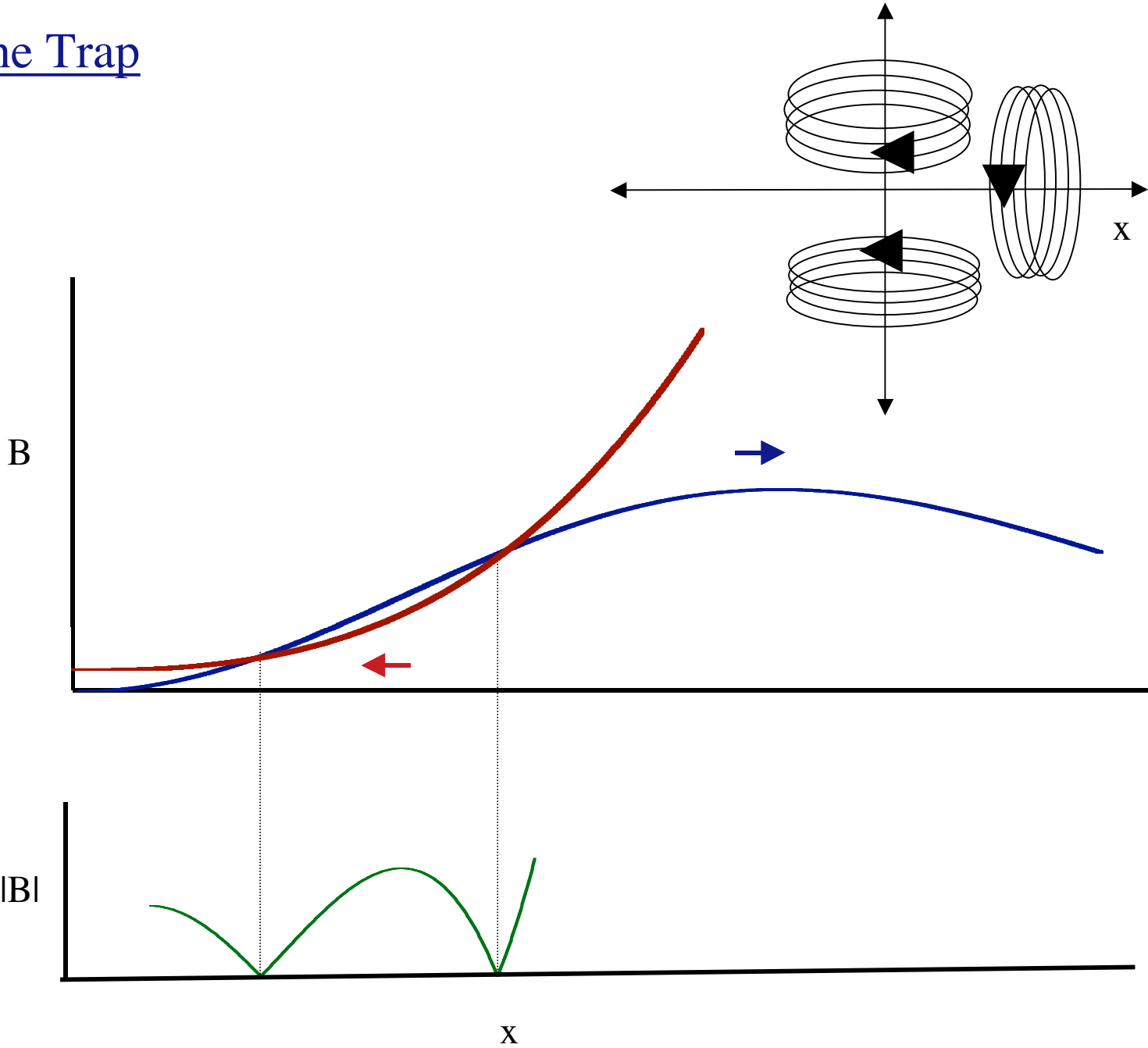
The trap



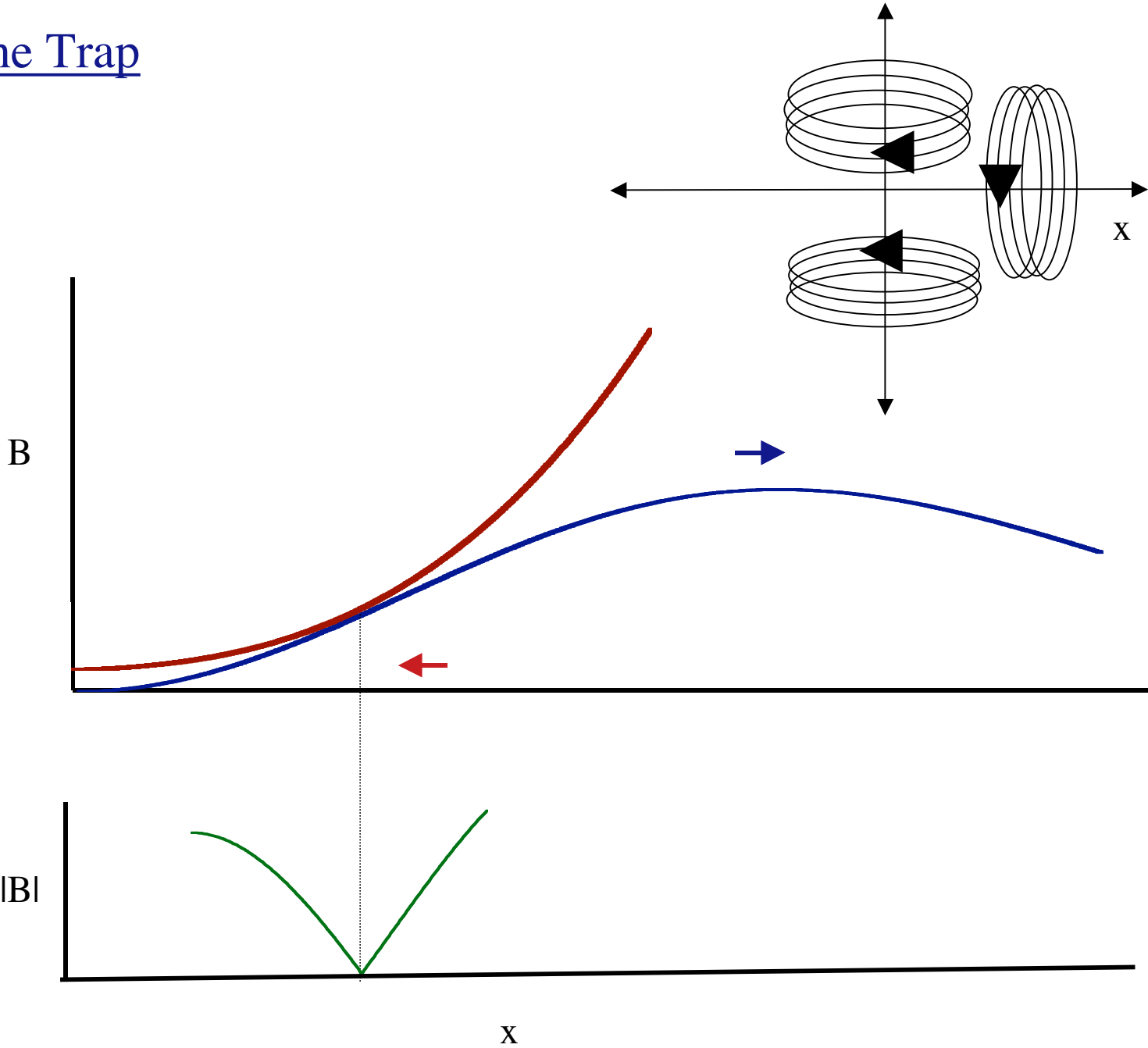
Forming the Trap



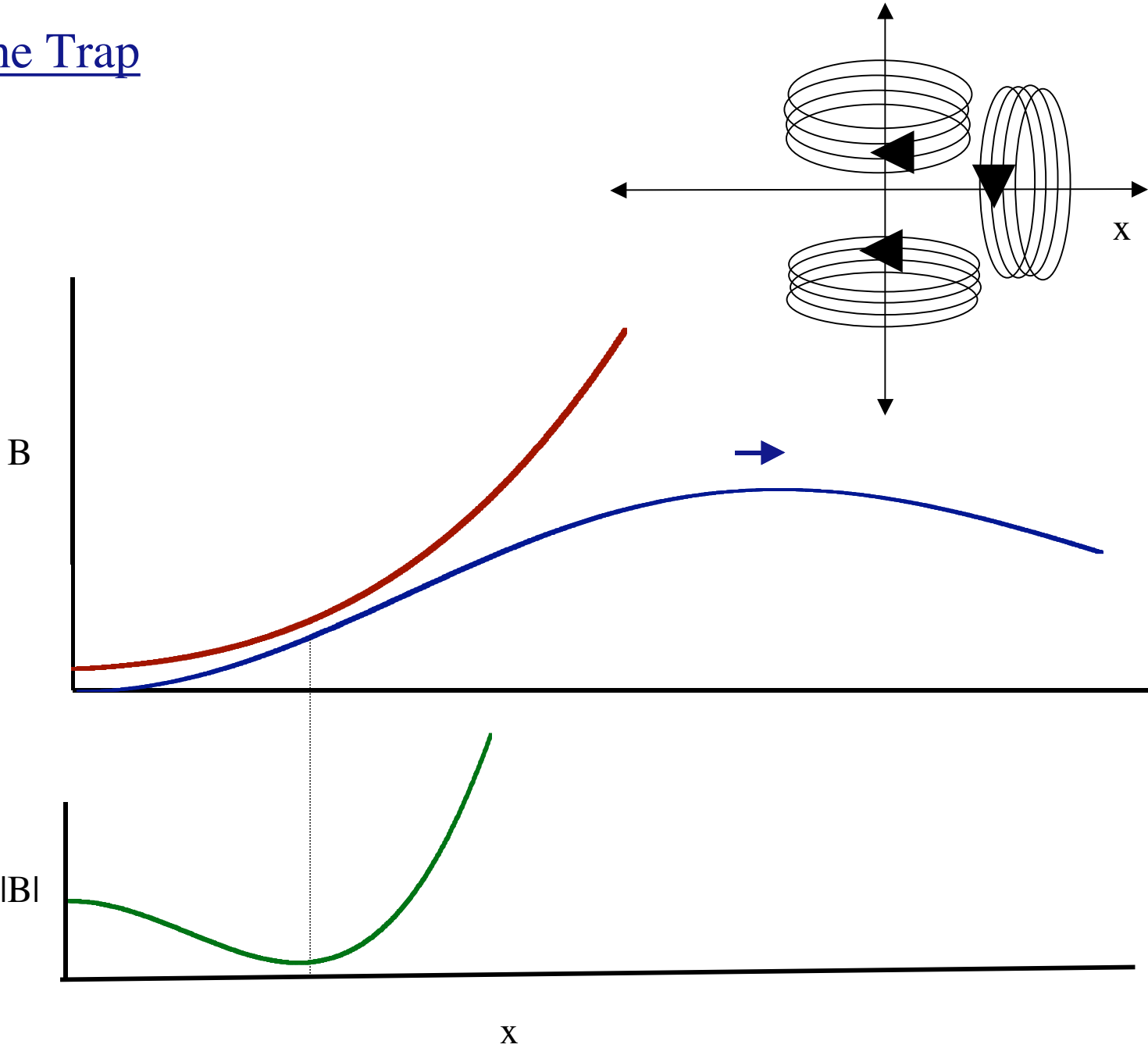
Forming the Trap



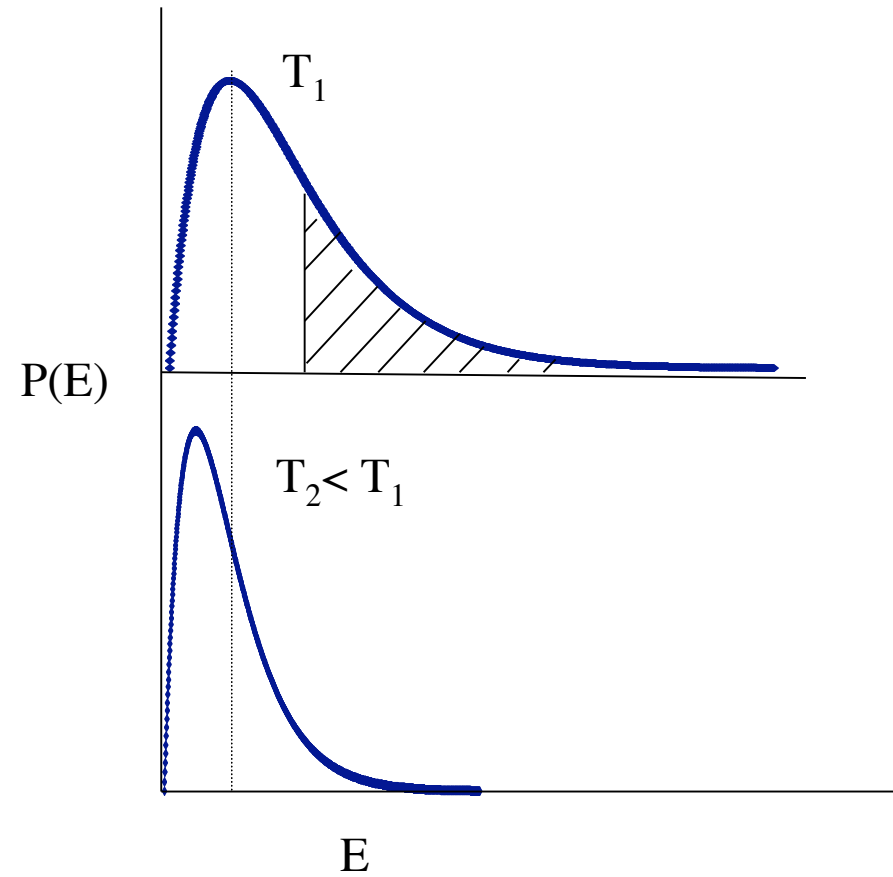
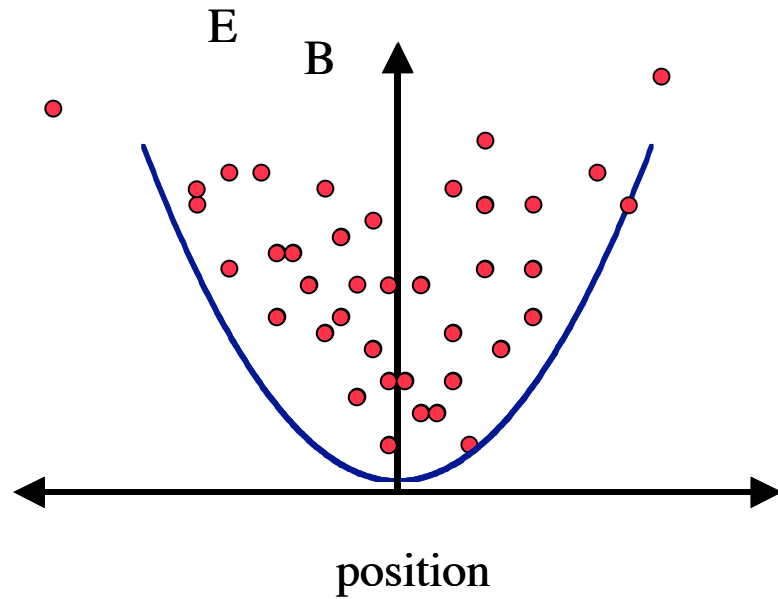
Forming the Trap



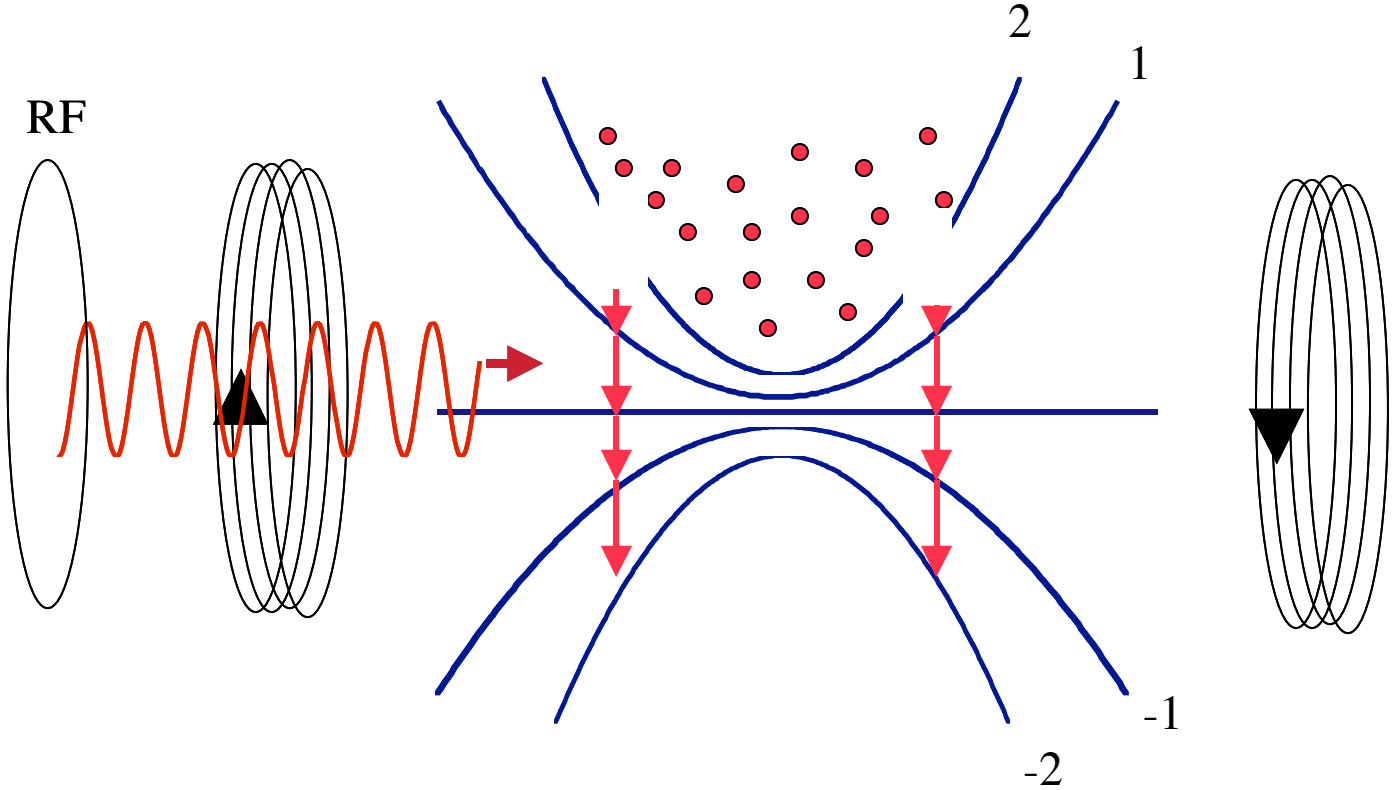
Forming the Trap



Evaporative Cooling: Truncating the Boltzmann Distribution



RF forced Evaporation



Elastic Collision Rate and Run Away Evaporation

$$\gamma_{el} = n\sigma_s V$$

$$n(\vec{r}) \propto N C \exp(-E/kT) \propto N C \exp[-U_{trap}(r)/kT]$$

$$C \iiint d^3r \exp[-\beta r^2/kT] = 1$$

$$C = \left(\frac{\beta}{kT}\right)^{\frac{3}{2}}$$

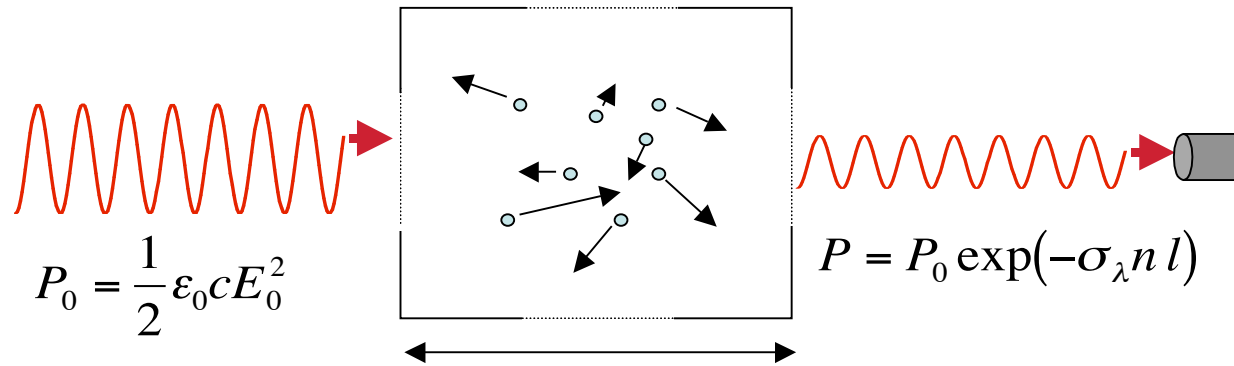
$$n(0) \propto NT^{-\frac{3}{2}}$$

$$V \propto T^{\frac{1}{2}}$$

$$\gamma_{el} = n\sigma_s V \propto \frac{N}{T}$$

$$\rho_{phase\ space} \propto NT^{-3}$$

Optical Depth and Runaway Evaporation



$$\sigma_\lambda n l = OD = \ln \frac{P_0}{P}$$

Following a similar procedure to that on the previous page we find

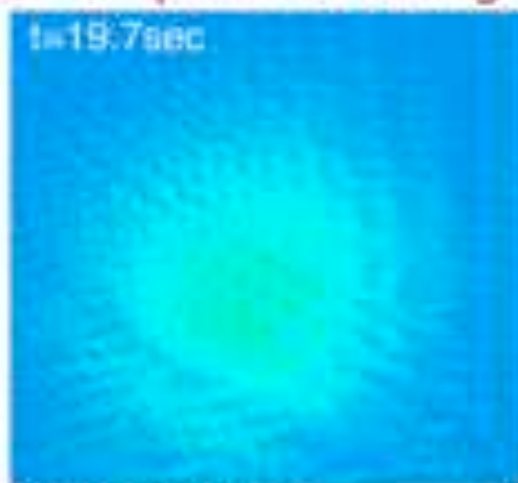
$$OD \propto \frac{N}{T} \propto \gamma_{el}$$

Linear Evaporative Cooling Ramp



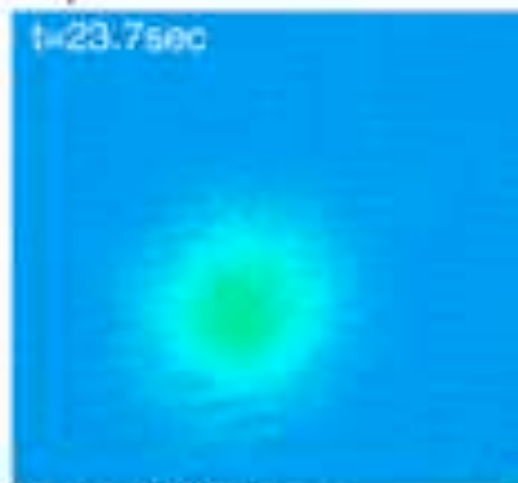
$t=8.3\text{sec}$

$N=4.9 \times 10^8$ atoms
 $T=190\mu\text{K}$ $D=1.3 \times 10^{-5}$



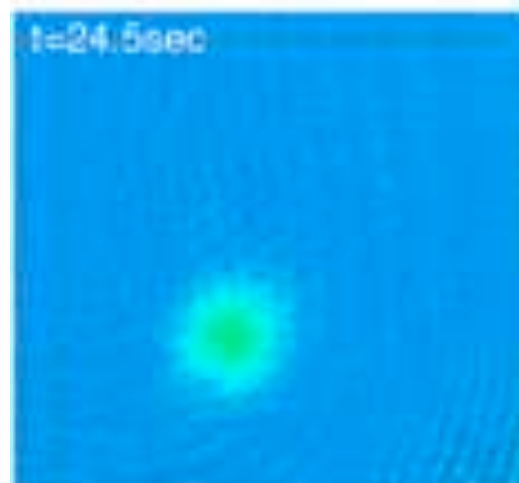
$t=19.7\text{sec}$

$N=2.8 \times 10^8$ atoms
 $T=85\mu\text{K}$ $D=8.7 \times 10^{-5}$



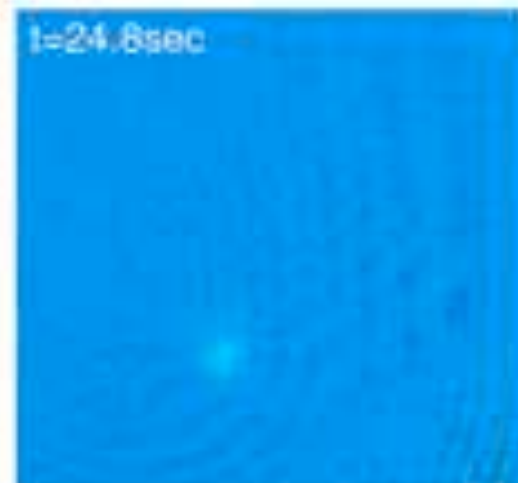
$t=23.7\text{sec}$

$N=9.8 \times 10^7$ atoms
 $T=25\mu\text{K}$ $D=1.2 \times 10^{-3}$



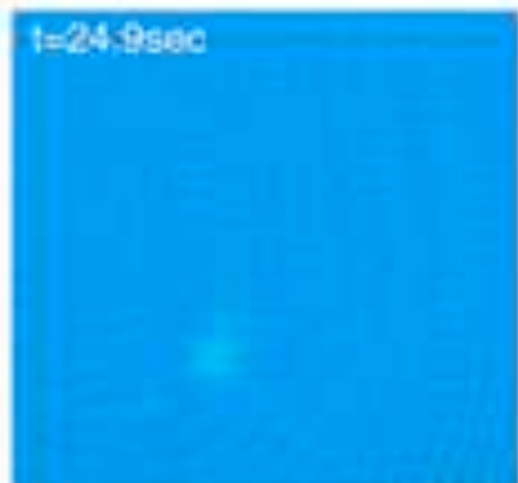
$t=24.5\text{sec}$

$N=4.3 \times 10^7$ atoms
 $T=8.3\mu\text{K}$ $D=0.014$



$t=24.8\text{sec}$

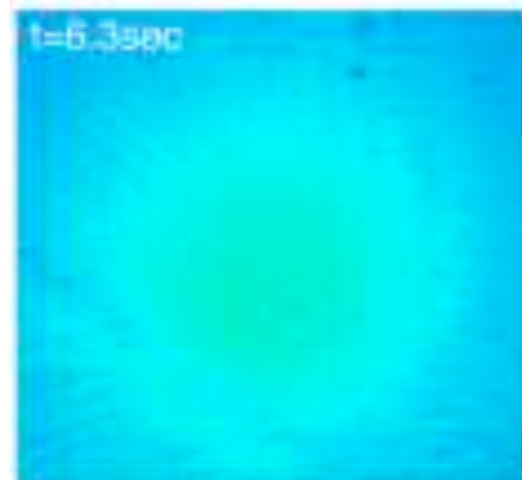
$N=3.7 \times 10^6$ atoms
 $T=2.3\mu\text{K}$ $D=0.058$



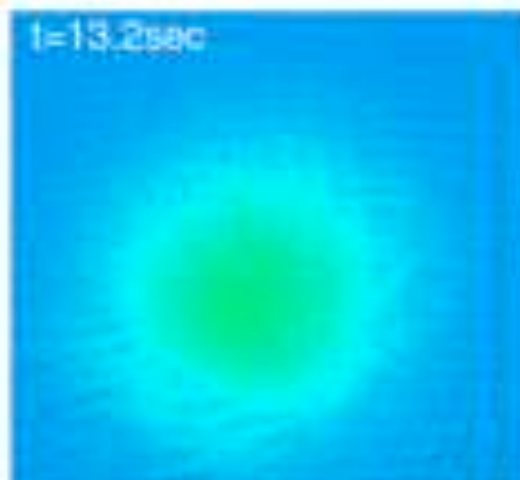
$t=24.9\text{sec}$

$N=1.0 \times 10^6$ atoms
 $T=2.4\mu\text{K}$ $D=0.014$

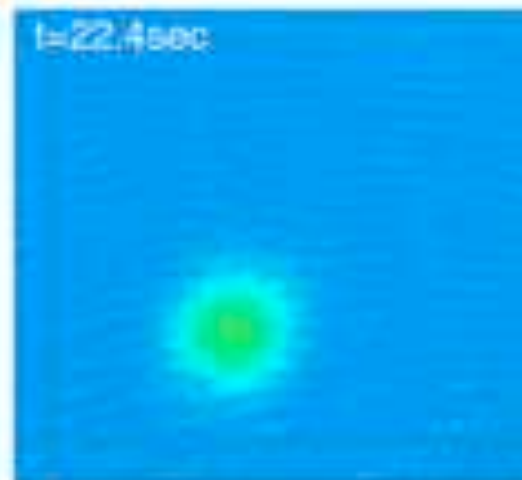
Logarithmic Evaporative Cooling Ramp



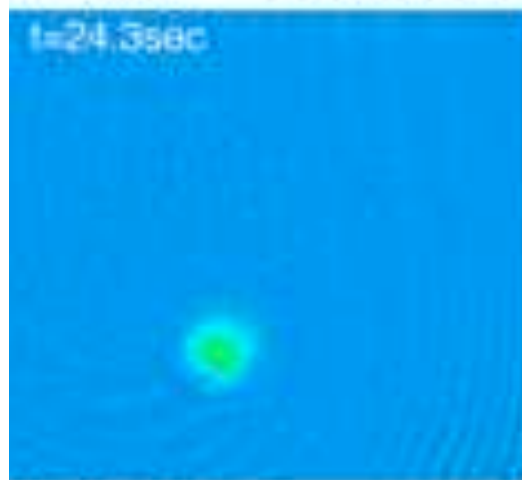
$t=6.3\text{sec}$
 $N=4.1 \times 10^8$ atoms
 $T=100\mu\text{K}$ $D=7.8 \times 10^{-5}$



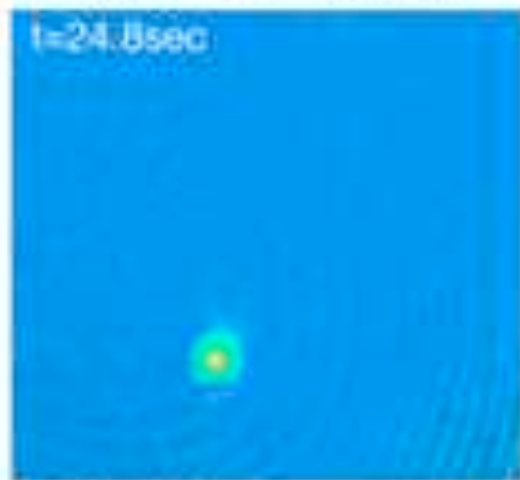
$t=13.2\text{sec}$
 $N=2.4 \times 10^8$ atoms
 $T=64\mu\text{K}$ $D=1.7 \times 10^{-4}$



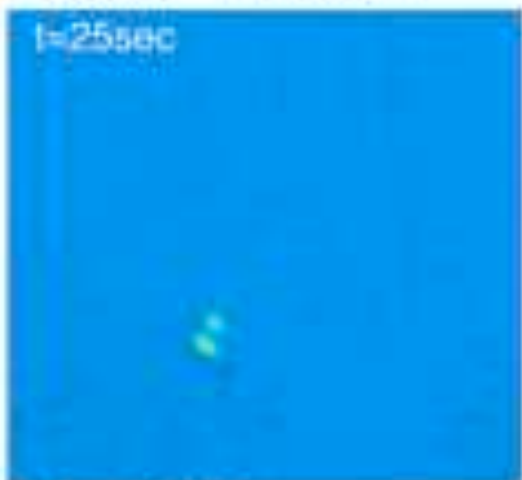
$t=22.4\text{sec}$
 $N=3.8 \times 10^7$ atoms
 $T=8.2\mu\text{K}$ $D=0.013$



$t=24.3\text{sec}$
 $N=1.3 \times 10^7$ atoms
 $T=2.2\mu\text{K}$ $D=0.23$



$t=24.8\text{sec}$
 $N=9.6 \times 10^6$ atoms
 $T=1.0\mu\text{K}$ $D=1.83$



$t=25\text{sec}$
 $N=1.4 \times 10^6$ atoms
 $T < 400\text{nK}$ $D \gg 2.6$

END OF LECTURE 2